Seismology and Seismic Hazard

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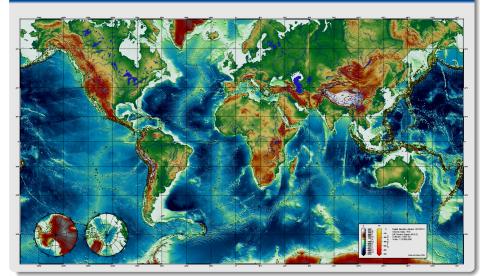
Seismology

- Comprises all about earthquakes and the propagation of seismic waves in the Earth.
- One of the main fields of solid-earth geophysics.
- Has provided the majority of our knowledge on Earth's interior.

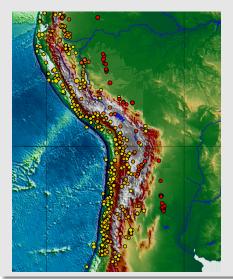
Seismics

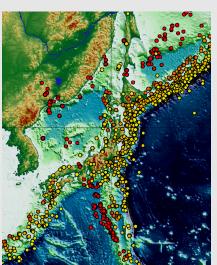
- Exploration of the deep and shallow subsurface with the help of artificial seismic waves.
- The perhaps most important field of applied geophysics.







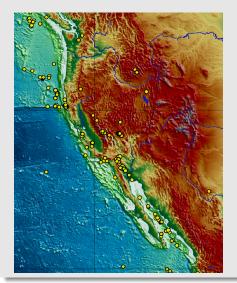






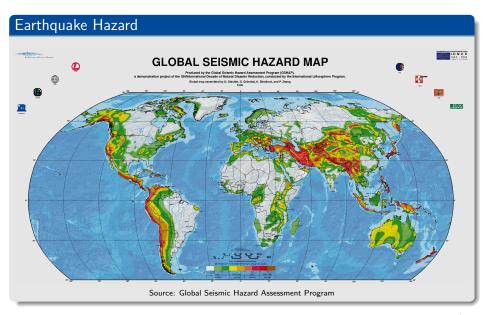






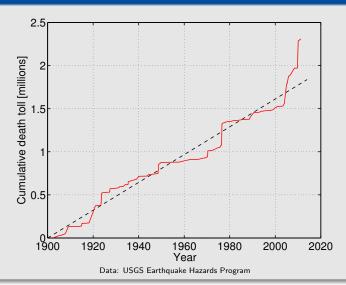








Earthquakes with 1000 or more Deaths 1900-2014





The First "Seismometer" (132 a.D.)





History of Seismology

1660	basic law of elasticity	R. Hooke
1821–22	differential equations of elasticity	C. Navier
		A. L. Cauchy
1830	theory of two fundamental types of elastic waves (P- and S-wave)	S. D. Poisson
1875	First "serious" seismometer	F. Gecchi
1887	theory of the first type of surface waves	J.W. Strutt (3.
		Lord Rayleigh)
1889	first recording of a distant earthquake	
1892	first compact seismometer, used at about	J. Milne
	40 stations	
1894	statistics of aftershocks	F. Omori
1903	12 degree scale for the intensity of earth-	G. Mercalli
	quakes based on the damage	



History of Seismology

1906–1913	detection of the liquid core of the earth and determination of its size	R. D. Oldham, B. Gutenberg
1909	detection of the crust-mantle discontinuity	A. Mohorovičić
1911	theory of a second type of surface waves	A. E. H. Love
1935	local magnitude as an "objective" measure of earthquake intensity	C. F. Richter
1936	detection of the inner, solid core	I. Lehmann
1954	frequency-magnitude relation of earthquakes	B. Gutenberg, C. F. Richter
1975	first successful short-term prediction of a strong earthquake	
1977	moment magnitude as a measure of earthquake source strength	H. Kanamori

Theory of Seismic Waves



The Navier-Cauchy Equations in Seismology

- Small, but spatially and temporally variable displacement $\vec{u}(\vec{x},t)$
- Neglect gravity
- Sign convention as in mathematics, physics, and engineering
- Elastic deformation



$$\rho \frac{\partial^2}{\partial t^2} \vec{u} = \begin{pmatrix} \frac{\partial \sigma_{11}}{\partial x_1} + \frac{\partial \sigma_{12}}{\partial x_2} + \frac{\partial \sigma_{13}}{\partial x_3} \\ \frac{\partial \sigma_{21}}{\partial x_1} + \frac{\partial \sigma_{22}}{\partial x_2} + \frac{\partial \sigma_{23}}{\partial x_3} \\ \frac{\partial \sigma_{31}}{\partial x_1} + \frac{\partial \sigma_{32}}{\partial x_2} + \frac{\partial \sigma_{33}}{\partial x_3} \end{pmatrix} = \operatorname{div}(\boldsymbol{\sigma})$$

(1)

with the stress tensor

$$\boldsymbol{\sigma} = \lambda \, \epsilon_{\mathbf{v}} \, \mathbf{1} + 2\mu \, \boldsymbol{\epsilon},$$

Theory of Seismic Waves



The Navier-Cauchy Equations in Seismology

the strain tensor ϵ consisting of the components

$$\epsilon_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right), \tag{3}$$

the volumetric strain

$$\epsilon_{v} = \epsilon_{11} + \epsilon_{22} + \epsilon_{33}, \tag{4}$$

the density ρ and the Lamé parameters of the medium λ and μ



(5)

(6)

The Navier-Cauchy Equations in 1D

Displacement u(x, t) instead of $\vec{u}(\vec{x}, t)$.



$$\frac{\partial^2}{\partial u^2}u =$$

$$\rho \frac{\partial^2}{\partial t^2} u(x,t) = \frac{\partial}{\partial x} \left((\lambda + 2\mu) \frac{\partial}{\partial x} u(x,t) \right) \quad \text{(1D wave equation)}$$

$$\sigma = (\lambda + 2\mu)\epsilon = (\lambda + 2\mu)\frac{\partial}{\partial x}u$$



Solution of the 1D Wave Equation

If λ and μ are constant:

$$\rho \frac{\partial^2}{\partial t^2} u(x,t) = (\lambda + 2\mu) \frac{\partial^2}{\partial x^2} u(x,t)$$
 (8)

Solution:

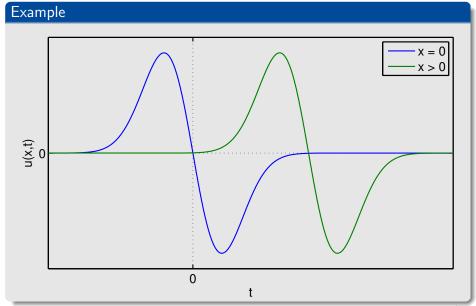
$$u(x,t) = f(t \pm sx) \tag{9}$$

where

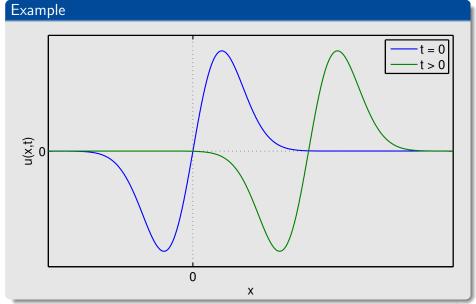
$$f=$$
 arbitrary function describing the shape of the wave $s=\sqrt{rac{
ho}{\lambda+2\mu}}=$ slowness

The wave moves in positive or negative x direction with a velocity $v = \frac{1}{s}$.











The Retarded Time

 $\tau = t \pm sx$ is called retarded time.

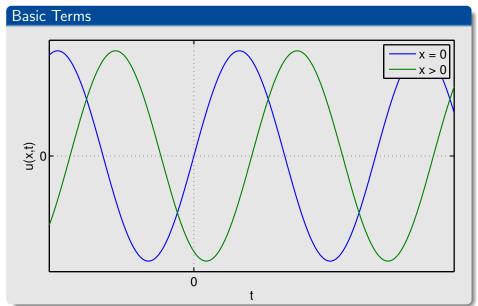
Meaning: At the position x and time t we observe what happened at the origin (x=0, e.g. earthquake focus) at the retarded time $\tau=t\pm sx$.

The Shape of the Wave

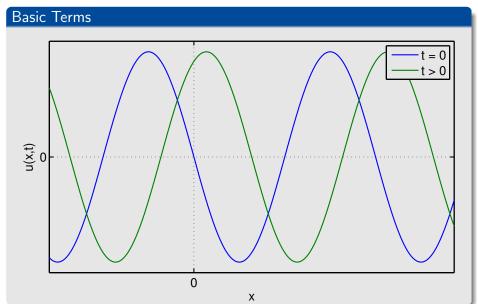
Examples for the function f:

- $f(\tau) = \begin{cases} 1 & \text{for } \tau \geq T \\ 0 & \text{else} \end{cases}$ describes a step-like shape (shock wave).
- $f(\tau) = \left\{ egin{array}{ll} 1 & {
 m for} \ |\tau| \leq rac{T}{2} \\ 0 & {
 m else} \end{array}
 ight.$ describes a boxcar-shaped wave.
- $f(\tau) = a \cos(\omega \tau)$, $f(\tau) = a \sin(\omega \tau)$ or $f(\tau) = a e^{i\omega \tau}$ describes a harmonic wave with an angular frequency ω and amplitude a.











Basic Terms

• Time domain:

Angular frequency: $\omega \left[\frac{1}{s}\right]$

Frequency: $\nu = \frac{\omega}{2\pi} \left[\frac{1}{s} \right]$

Period: $T = \frac{1}{\nu} = \frac{2\pi}{\omega}$ [s]

• Spatial wave pattern:

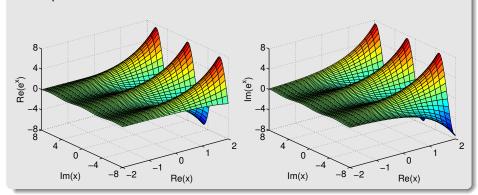
Wave number: $k = \omega s \left[\frac{1}{m}\right]$

Wavelength: $L = \frac{2\pi}{k} = \frac{1}{\nu s}$ [m]



The Complex Exponential Function vs. Sine and Cosine

With $e^{i\phi}=\cos\phi+i\sin\phi$, the complex exponential function combines the real exponential function with the sine and cosine functions.





(10)

The Complex Exponential Function vs. Sine and Cosine



 $\operatorname{Re}\left(e^{i\omega\tau}\right) = \cos\left(\omega\tau\right) \quad \text{and} \quad \operatorname{Im}\left(e^{i\omega\tau}\right) = \sin\left(\omega\tau\right)$

Real part and imaginary part of the complex solution can be considered as independent real solutions.

Derivatives of the complex solutions are simpler than those of the real solutions:

$$\frac{\partial}{\partial \tau} e^{i\omega\tau} = i\omega e^{i\omega\tau} \tag{11}$$

while

$$\frac{\partial}{\partial \tau}\cos(\omega\tau) = -\omega\sin(\omega\tau)$$
 and $\frac{\partial}{\partial \tau}\sin(\omega\tau) = \omega\cos(\omega\tau)$ (12)



Seismic Waves in 3D

Complications towards the 1D case:

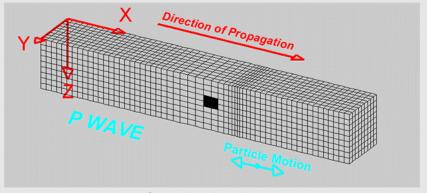
- Displacement $\vec{u}(\vec{x}, t)$ is a vector.
- Propagation in arbitrary direction in space instead of the positive or negative x axis only.



Fundamental Types of Body Waves

Two types of independent plane waves in an infinite, homogeneous elastic medium:

Compressional wave (longitudinal wave, primary wave)

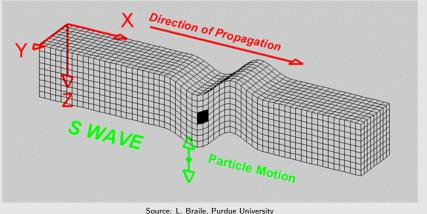


Source: L. Braile, Purdue University



Fundamental Types of Body Waves

• Shear wave (transverse wave, secondary wave)





Plane Waves in Infinite, Homogeneous, and Isotropic Media

Plane wave: $\vec{u}(\vec{x}, t)$ is constant on parallel planes.

Mathematical description:

$$\vec{u}(\vec{x},t) = f(t-\vec{s}\cdot\vec{x})\vec{a}$$
 (13)

or for a harmonic wave:

$$\vec{u}(\vec{x},t) = e^{i\omega(t-\vec{s}\cdot\vec{x})}\vec{a}$$
 (14)

where

$$\vec{s}$$
 = slowness vector
 \vec{a} = amplitude vector (constant)

The wave moves in direction of \vec{s} with a velocity $v = \frac{1}{|\vec{s}|}$.



Plane Waves in Infinite, Homogeneous, and Isotropic Media

 $\boldsymbol{\sigma} = -s f'(t - sx_1) \begin{pmatrix} (\lambda + 2\mu)a_1 & \mu a_2 & \mu a_3 \\ \mu a_2 & \lambda a_1 & 0 \\ \mu a_3 & 0 & \lambda a_1 \end{pmatrix}$

$$\epsilon = -s f'(t - sx_1) \begin{pmatrix} a_1 & \frac{1}{2}a_2 & \frac{1}{2}a_3 \\ \frac{1}{2}a_2 & 0 & 0 \\ \frac{1}{2}a_3 & 0 & 0 \end{pmatrix}$$

(17)

(15)

$$\vec{u}(\vec{x},t) = f(t-sx_1)\vec{a}$$

Simplest version: propagation in
$$x_1$$
 direction, $\vec{s} = \begin{pmatrix} s \\ 0 \\ 0 \end{pmatrix}$



(18)

Plane Waves in Infinite, Homogeneous, and Isotropic Media

If λ and μ are constant:

div
$$(\sigma) = s^2 f''(t - sx_1) \begin{pmatrix} (\lambda + 2\mu)a_1 \\ \mu a_2 \\ \mu a_3 \end{pmatrix}$$

Insert into the Navier-Cauchy equations:

$$\rho \frac{\partial^2}{\partial t^2} \vec{u} = \rho f''(t - sx_1) \vec{a}$$

$$= \operatorname{div}(\boldsymbol{\sigma}) = s^2 f''(t - sx_1) \begin{pmatrix} (\lambda + 2\mu)a_1 \\ \mu a_2 \\ \mu a_3 \end{pmatrix}$$
(20)

Can only be satisfied if $a_2=a_3=0$ and $s=\sqrt{\frac{\rho}{\lambda+2\mu}}$ (longitudinal polarization) or $a_1=0$ and $s=\sqrt{\frac{\rho}{\mu}}$ (transverse polarization).



Plane Waves in Infinite, Homogeneous, and Isotropic Media

General case (assignment 2): Navier-Cauchy equations can be satisfied only if either \vec{a} is parallel (or opposite) to \vec{s} or normal to \vec{s} .

Transverse wave: \vec{a} is normal to \vec{s} ($\vec{a} \cdot \vec{s} = 0$)

$$|\vec{s}|^2 = \frac{\rho}{\mu}, \quad v_s = \frac{1}{|\vec{s}|} = \sqrt{\frac{\mu}{\rho}}$$
 (21)

Longitudinal wave: \vec{a} is parallel or opposite to \vec{s}

$$|\vec{s}|^2 = \frac{\rho}{\lambda + 2\mu}, \quad v_p = \frac{1}{|\vec{s}|} = \sqrt{\frac{\lambda + 2\mu}{\rho}}$$
 (22)



Comparison with Sound Waves in Liquids and Gases

The longitudinal wave is similar to sound waves in liquids and gases, while the transverse wave has no counterpart in liquids and gases.

Seismic Velocities

Medium	Longitudinal wave $\left[\frac{km}{s}\right]$	Transverse wave $\left[\frac{km}{s}\right]$
air	0.34	-
water	1.45	-
wood	about 3	about 1.8
Earth*	5.8–13.7	3.4–7.2

^{*}Parametric Earth Models (PEM), not valid for the shallow subsurface



Seismic Velocities

$$\frac{v_p}{v_s} = \sqrt{\frac{\lambda + 2\mu}{\mu}} \ge \sqrt{\frac{4}{3}} \approx 1.15 \tag{23}$$



Velocity v_p of the longitudinal wave is always higher than the velocity v_s of the transverse wave.



Longitudinal wave always arrives prior to the transverse wave.



longitudinal wave = primary wave (P-wave) transverse wave = secondary wave (S-wave)



Typical v_p - v_s Ratios

For solid rocks:

$$\frac{v_p}{v_s} = \sqrt{\frac{\lambda + 2\mu}{\mu}} \approx \sqrt{3} \approx 1.7 \tag{24}$$

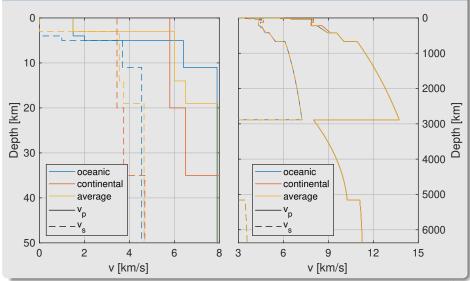
for $\lambda \approx \mu$.

• For soil or unconsolidated rocks:

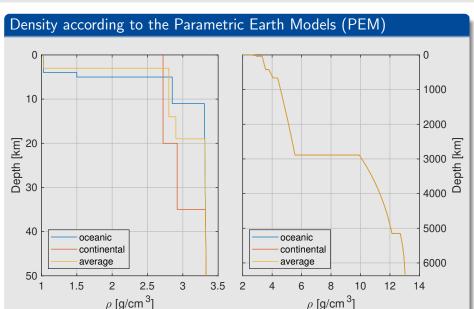
$$\frac{v_p}{v_s} \approx 2.5 \tag{25}$$



Seismic Velocities according to the Parametric Earth Models (PEM)

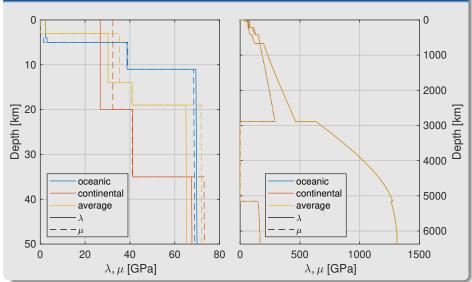








Lamé Parameters according to the Parametric Earth Models (PEM)



Seismic Body Waves



Typical P-wave Velocities in the Shallow Subsurface

Medium	$V_p\left[\frac{\mathrm{km}}{\mathrm{s}}\right]$
weathering zone	0.1–0.5
dry sand	0.3-0.6
water-saturated sand	1.3-1.8
sandstone	1.8-4
pit coal	1.6-1.9

	_
Medium	$v_p\left[\frac{\mathrm{km}}{\mathrm{s}}\right]$
clay	1.2-2.8
claystone	2.2-4.2
limestone	3–6
halite	4.5–6.5
granite	5–6.5



The Point-Force Solution

- Infinite, homogeneous medium with parameters ρ , λ , and μ (like plane wave consideration).
- Assume that a given force $\vec{F}(t)$ acts at the origin $(\vec{x} = \vec{0})$.

Respective solution of the Navier-Cauchy equations:

$$\vec{u}_{f}(\vec{x},t) = \frac{s_{p}^{2}}{4\pi\rho r} \mathbf{P} \vec{F}(t - s_{p}r) + \frac{s_{s}^{2}}{4\pi\rho r} (\mathbf{1} - \mathbf{P}) \vec{F}(t - s_{s}r)$$
(26)
+
$$\frac{1}{4\pi\rho r^{3}} (3\mathbf{P} - \mathbf{1}) \int_{s_{p}r}^{s_{s}r} \tau \vec{F}(t - \tau) d\tau$$
(27)

where

$$r = |\bar{x}|$$

$$\mathbf{P} = \vec{e} \, \vec{e}^T = \text{projection on radial direction}, \quad \vec{e} = \frac{\vec{x}}{r}$$



The Point-Force Solution

$$\vec{u}_{f}(\vec{x},t) = \frac{s_{p}^{2}}{4\pi\rho r} \mathbf{P} \vec{F}(t - s_{p}r) + \frac{s_{s}^{2}}{4\pi\rho r} (\mathbf{1} - \mathbf{P}) \vec{F}(t - s_{s}r)$$
(28)

$$+ \frac{1}{4\pi\rho r^{3}} (3\mathbf{P} - \mathbf{1}) \int_{s_{p}r}^{s_{s}r} \tau \vec{F}(t - \tau) d\tau$$
(29)

Example:



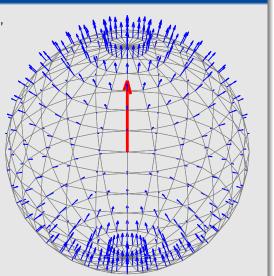


The Point-Force Solution

Spatial pattern of the first term,

P₽,

$$\vec{F} = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$



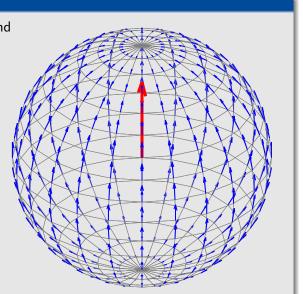


The Point-Force Solution

Spatial pattern of the second term,

$$(\mathbf{1} - \mathbf{P}) \vec{F}$$
,

$$\vec{F} = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$





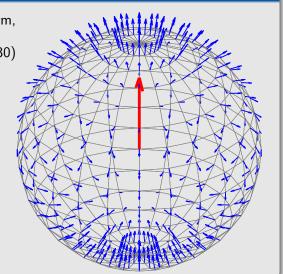
The Point-Force Solution

Spatial pattern of the third term,

$$(3P - 1)\vec{F}$$
,

(30)

$$\vec{F} = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

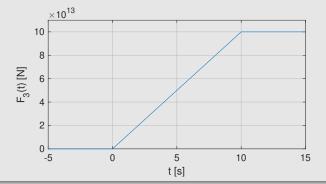




The Point-Force Solution

Example: second term (S-wave) for

$$\vec{F}(t) = \begin{pmatrix} 0 \\ 0 \\ F_3(t) \end{pmatrix}$$





Force Couples

Solution for a single point force causes an overall displacement in direction of the force.



not possible



Consider a couple of opposite forces \vec{F} and $-\vec{F}$ displaced by a small vector \vec{a} (at $\frac{\vec{a}}{2}$ and $-\frac{\vec{a}}{2}$).



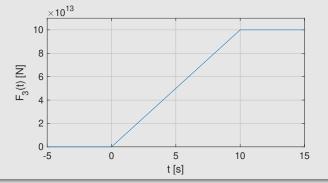
$$\vec{u}(\vec{x},t) = \vec{u}_f(\vec{x} - \frac{\vec{a}}{2},t) - \vec{u}_f(\vec{x} + \frac{\vec{a}}{2},t)$$



The Point-Force Solution

Example: second term (S-wave) for

$$\vec{F}(t) = \begin{pmatrix} 0 \\ 0 \\ F_3(t) \end{pmatrix}, \quad \vec{a} = \begin{pmatrix} 100 \text{ m} \\ 0 \\ 0 \end{pmatrix}$$





The Seismic Moment Tensor

<u>V</u>

Approximation in the limit $|\vec{a}| \rightarrow 0$:

$$\vec{u}(\vec{x},t) \approx -\nabla \vec{u}_{f}(\vec{x},t)\vec{a}$$

$$= -\text{div}\left(\frac{s_{p}^{2}}{4\pi\rho r}\text{PM}(t-s_{p}r) + \frac{s_{s}^{2}}{4\pi\rho r}(\mathbf{1}-\text{P})\text{M}(t-s_{s}r)\right)$$

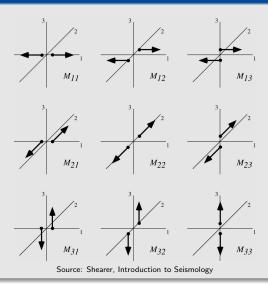
$$+\frac{1}{4\pi\rho r^{3}}(3\text{P}-\mathbf{1})\int_{s_{p}r}^{s_{s}r} \tau \text{M}(t-\tau)d\tau$$
(33)

with the seismic moment tensor (centroid moment tensor, CMT)

$$\mathbf{M}(t) = \vec{F}(t)\vec{a}^T \quad [Nm] \tag{34}$$



Components of the Seismic Moment Tensor





The Permanent Displacement



Perform the derivatives (div) in Eq. 33.



Terms proportional to \mathbf{M} and terms proportional to $\mathbf{M} = \frac{d}{dt}\mathbf{M}$

Terms proportional to M:

 Persist after the earthquake has terminated, i. e., after M has become constant.



Describe the permanent displacement caused by the earthquake.

• Decrease like $\frac{1}{r^2}$ or faster.



The Transient Displacement

Terms proportional to **M**:

- - Arise from applying the chain rule to terms of the form $\frac{\partial}{\partial x} \mathbf{M}(t sr)$.
 - Vanish after the earthquake has terminated.



Seismic waves

$$\vec{u}(\vec{x},t) = \frac{s_p^3}{4\pi\rho r} \mathbf{P} \dot{\mathbf{M}}(t - s_p r) \vec{e} + \frac{s_s^3}{4\pi\rho r} (\mathbf{1} - \mathbf{P}) \dot{\mathbf{M}}(t - s_s r) \vec{e} \qquad (35)$$

$$\uparrow \qquad \qquad \uparrow$$
P-wave S-wave

with
$$\vec{e} = \frac{\vec{x}}{r}$$
 and $\mathbf{P} = \vec{e} \vec{e}^T$

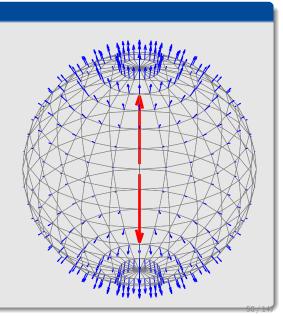


Seismic Wave Patterns

P-wave radiation pattern

 $\dot{\mathbf{PMe}}$

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$



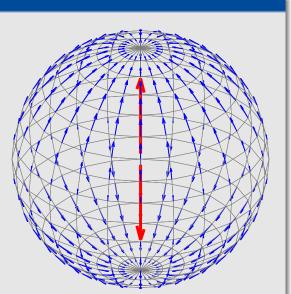


Seismic Wave Patterns

S-wave radiation pattern

$$(1 - P) \dot{M} \vec{e}$$

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$



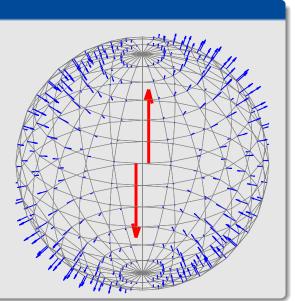


Seismic Wave Patterns

P-wave radiation pattern

 \overrightarrow{PMe}

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$$



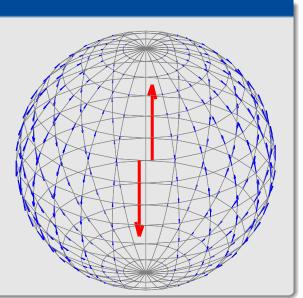


Seismic Wave Patterns

S-wave radiation pattern

$$(1 - P) \dot{M} \vec{e}$$

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$$





Symmetry of the Seismic Moment Tensor

Nondiagonal components of ${\bf M}$ cause an overall rotation.



not possible



M must be symmetric: $M_{ji} = M_{ij}$.

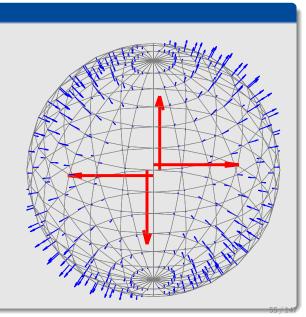


Seismic Wave Patterns

P-wave radiation pattern

 \overrightarrow{PMe}

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$$



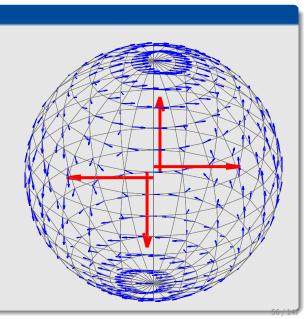


Seismic Wave Patterns

S-wave radiation pattern

$$(1 - P) \dot{M} \vec{e}$$

$$\dot{\mathbf{M}} = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$$



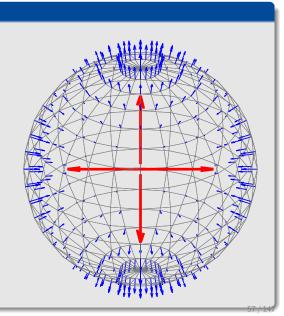


Seismic Wave Patterns

P-wave radiation pattern

 $\dot{\mathbf{PMe}}$

$$\dot{\mathbf{M}} = \begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$



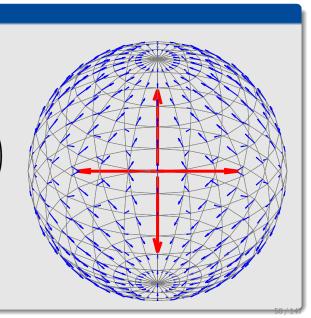


Seismic Wave Patterns

S-wave radiation pattern

$$(1 - P) \dot{M} \vec{e}$$

$$\dot{\mathbf{M}} = \begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$





The Scalar Seismic Moment

lf

$$\mathbf{M} = \begin{pmatrix} 0 & 0 & M_0 \\ 0 & 0 & 0 \\ M_0 & 0 & 0 \end{pmatrix} \quad \text{or} \quad \mathbf{M} = \begin{pmatrix} -M_0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & M_0 \end{pmatrix}$$
(36)

(or similar), M_0 is called (scalar) seismic moment.

In general:

$$M_0 = \frac{M_1 - M_3}{2} \tag{37}$$

where M_1 , M_2 , and M_3 are the eigenvalues of \mathbf{M} in descending order.



The Scalar Seismic Moment

Alternative interpretation of the seismic moment:

$$M_0 = \mu A \overline{u} \tag{38}$$

where

 $A = \text{size of the rupture area } [\text{m}^2]$

 $\overline{u} = \mathsf{mean} \; \mathsf{displacement} \; \mathsf{along} \; \mathsf{the} \; \mathsf{rupture} \; \mathsf{area} \; [\mathsf{m}]$





Amplitudes of Body Waves

P-wave displacement:

$$\vec{u}_p(\vec{x},t) = \frac{s_p^3}{4\pi\rho r} \mathbf{P} \dot{\mathbf{M}}(t - s_p r) \vec{e}$$
 (39)

Maximum displacement occurs in the directions of the first and third principal axes of $\dot{\mathbf{M}}$:

$$|\vec{u}_p|_{\text{max}} = \frac{s_p^3}{4\pi\rho r} |\dot{M}_0|_{\text{max}} \tag{40}$$



Amplitudes of Body Waves

S-wave displacement:

$$\vec{u}_s(\vec{x},t) = \frac{s_s^3}{4\pi\rho r} (\mathbf{1} - \mathbf{P}) \dot{\mathbf{M}} (t - s_s r) \vec{e}$$
 (41)

Maximum displacement occurs in the directions 45° between the first and third principal axis of $\dot{\mathbf{M}}$:

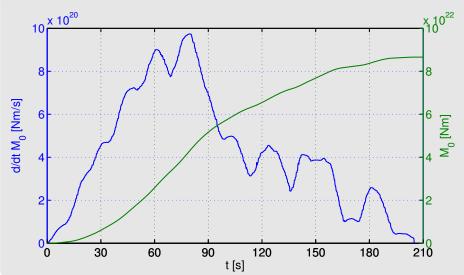
$$|\vec{u}_s|_{\text{max}} = \frac{s_s^3}{4\pi\rho r} |\dot{M}_0|_{\text{max}}$$
 (42)



$$\frac{|\vec{u}_s|_{\text{max}}}{|\vec{u}_p|_{\text{max}}} = \frac{s_s^3}{s_p^3} = \left(\frac{v_p}{v_s}\right)^3 \approx 5 \tag{43}$$



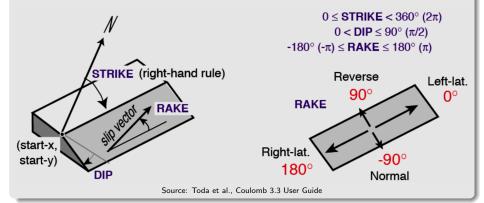






The Seismic Moment Tensor for a Double Force Couple

Definition of strike ϕ , dip δ , and rake λ according to Aki and Richards (1980)





The Seismic Moment Tensor for a Double Force Couple

Step 1: Start with a force couple in x_2 direction displaced in x_3 direction:

$$\vec{F} = \begin{pmatrix} 0 \\ F \\ 0 \end{pmatrix}, \quad \vec{a} = \begin{pmatrix} 0 \\ 0 \\ a \end{pmatrix} \tag{44}$$

Step 2: Rotate \vec{F} and \vec{a} counterclockwise by the rake angle λ in the x_1 - x_2 plane:

$$\vec{F}_{\lambda} = \mathbf{R}_{\lambda}\vec{F}, \quad \vec{a}_{\lambda} = \mathbf{R}_{\lambda}\vec{a} = \vec{a}$$
 (45)

with

$$\mathbf{R}_{\lambda} = \begin{pmatrix} \cos \lambda & -\sin \lambda & 0\\ \sin \lambda & \cos \lambda & 0\\ 0 & 0 & 1 \end{pmatrix} \tag{46}$$



The Seismic Moment Tensor for a Double Force Couple

Step 3: Rotate \vec{F}_{λ} and \vec{a}_{λ} clockwise by the dip angle δ in the x_1 - x_3 plane:

$$\vec{F}_{\lambda\delta} = \mathbf{R}_{\delta}\vec{F}_{\lambda} = \mathbf{R}_{\delta}\mathbf{R}_{\lambda}\vec{F}, \quad \vec{a}_{\lambda\delta} = \mathbf{R}_{\delta}\vec{a}_{\lambda} = \mathbf{R}_{\delta}\mathbf{R}_{\lambda}\vec{a}$$
 (47)

with

$$\mathbf{R}_{\delta} = \begin{pmatrix} \cos \delta & 0 & \sin \delta \\ 0 & 1 & 0 \\ -\sin \delta & 0 & \cos \delta \end{pmatrix} \tag{48}$$



The Seismic Moment Tensor for a Double Force Couple

Step 4: Rotate $\vec{F}_{\lambda\delta}$ and $\vec{a}_{\lambda\delta}$ clockwise by the strike angle ϕ in the x_1 - x_2 plane:

$$\vec{F}_{\lambda\delta\phi} = \mathbf{R}_{\phi}\vec{F}_{\lambda\delta} = \mathbf{R}_{\phi}\mathbf{R}_{\delta}\mathbf{R}_{\lambda}\vec{F}, \quad \vec{a}_{\lambda\delta\phi} = \mathbf{R}_{\phi}\vec{a}_{\lambda\delta} = \mathbf{R}_{\phi}\mathbf{R}_{\delta}\mathbf{R}_{\lambda}\vec{a} \tag{49}$$

with

$$\mathbf{R}_{\phi} = \begin{pmatrix} \cos \phi & \sin \phi & 0 \\ -\sin \phi & \cos \phi & 0 \\ 0 & 0 & 1 \end{pmatrix}$$
 (50)



$$\vec{F}_{\lambda\delta\phi} = \mathbf{R}\vec{F}, \quad \vec{a}_{\lambda\delta\phi} = \mathbf{R}\vec{a} \quad \text{with} \quad \mathbf{R} = \mathbf{R}_{\phi}\mathbf{R}_{\delta}\mathbf{R}_{\lambda}$$
 (51)



The Seismic Moment Tensor for a Double Force Couple



$$\mathbf{M} = \vec{F}_{\lambda\delta\phi} \vec{a}_{\lambda\delta\phi}^{T} + \vec{a}_{\lambda\delta\phi} \vec{F}_{\lambda\delta\phi}^{T} = \left(\mathbf{R}\vec{F}\right) \left(\mathbf{R}\vec{a}\right)^{T} + \left(\mathbf{R}\vec{a}\right) \left(\mathbf{R}\vec{F}\right)^{T}$$
(52)

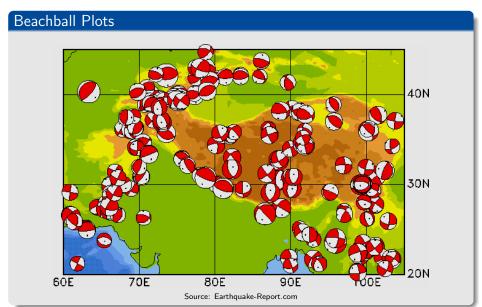
$$= \mathbf{R} \left(\vec{F} \vec{a}^T + \vec{a} \vec{F}^T \right) \mathbf{R}^T \tag{53}$$

$$= \mathbf{R} \left(\begin{pmatrix} 0 \\ F \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ 0 \\ a \end{pmatrix}^T + \begin{pmatrix} 0 \\ 0 \\ a \end{pmatrix} \begin{pmatrix} 0 \\ F \\ 0 \end{pmatrix}^T \right) \mathbf{R}^T$$
 (54)

$$= M_0 \mathbf{R} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} \mathbf{R}^T \tag{55}$$

with $M_0 = Fa$







Beachball Plots

Basic assumption: $\dot{\mathbf{M}}(t)$ has the same shape as $\mathbf{M}(t)$,

$$\mathbf{M}(t) = f(t)\mathbf{M}, \quad \dot{\mathbf{M}}(t) = \dot{f}(t)\mathbf{M}$$
 (56)

where **M** is the total seismic moment, and f(t) increases from 0 to 1.

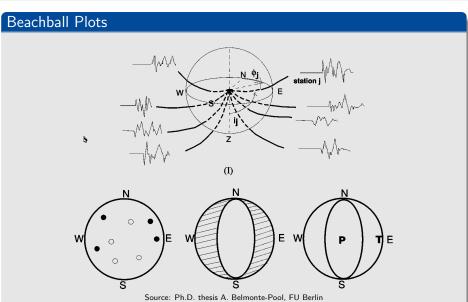
P-wave radiation pattern:

$$\vec{u} \propto \mathbf{PM}\vec{e} = ((\mathbf{M}\vec{e}) \cdot \vec{e})\vec{e}$$
 (57)



P-wave arrives with compression first if $(\mathbf{M}\vec{e})\cdot\vec{e}>0$ and with dilatation if $(\mathbf{M}\vec{e})\cdot\vec{e}<0$.

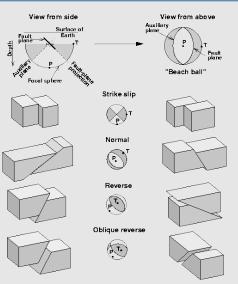






Beachball Plots

- Directions where the P-wave arrives first with compression $((\mathbf{M}\vec{e})\cdot\vec{e}>0)$ are colored
- Directions where the P-wave arrives first with dilatation $(\mathbf{M}\vec{e})\cdot\vec{e}<0$ are left white.
- Projection of the lower half of the sphere is plotted (sometimes stereographic projection, but mostly equal-area projection).



Source: US Geological Survey

Earthquake Source Theory



Beachball Plots

If the eigenvalues of \mathbf{M} are M_0 , 0, and $-M_0$, the sphere consists of 4 equal quadrants.

Examples of beachball plots:

- \bullet normal fault ($\lambda = -90^\circ)$ for different dip angles δ
- ullet reverse fault $(\lambda=90^\circ)$ for different dip angles δ
- ullet transform fault $(\lambda=0^\circ)$ for different dip angles δ
- ullet fault dipping at $\delta=45^\circ$ for different rake angles λ
- fault dipping at $\delta=45^\circ$ for different rake angles λ with additional isotropic expansion



Intensity and Magnitude

Intensity describes the severity of an earthquake in terms of its effects on the Earth's surface and on humans and their structures.

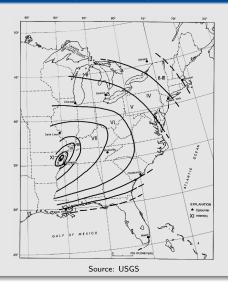
- Usually written as a Roman numeral.
- Goes back to a 12 level scale (originally 10) from I (not felt) to XII (total destruction) named after G. Mercalli (1850–1914).
- Several extensions / refinements: MCS (Mercalli-Cancani-Sieberg) scale, MWN (Mercalli-Wood-Neumann) scale, MSK scale (Medvedev, Sponheuer & Karnik, 1964), EMS-98 scale (European Macroseismic Scale, 2000).

Magnitude characterizes the size of an earthquake using measured values.

- Usually written as an Arabic numeral with one decimal digit.
- Several different magnitude definitions.
- Logarithmic scale.



Example of an Isoseismal Map of Earthquake Intensity





General Definition of Earthquake Magnitude

If X is any physically measured property of an earthquake, e. g.

- ullet total seismic moment M_0 or
- maximum ground displacement $|\vec{u}|_{\mathsf{max}}$,

the corresponding earthquake magnitude is defined by

$$M_X = e \log_{10} \left(\frac{X}{X_0} \right) \tag{58}$$

where

- X_0 = measured value for an earthquake of M_X = 0 under the same conditions
 - e= factor used for making different magnitude definitions consistent (mostly e=1)



General Definition of Earthquake Magnitude

If X is a property related to any point different from the earthquake focus, X_0 is a function of distance Δ and depth h (and other properties).



$$M_X = e \log_{10} \left(\frac{X}{X_0(\Delta, h)} \right) = e \log_{10} X + \sigma(\Delta, h)$$
 (59)

with the distance-depth correction function

$$\sigma(\Delta, h) = -e \log_{10} X_0(\Delta, h)$$
 (60)

This only makes sense if the distance-depth dependence of X is independent of X itself.



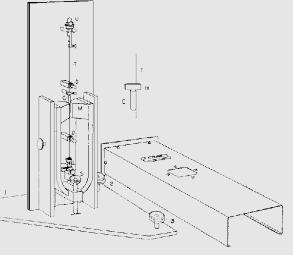
Upper und Lower Limits of Magnitude Scales

- All magnitude scales are from their definition open and both ends.
- Upper limits on Earth are introduced by geological constraints and by the process of wave propagation.
- Negative magnitudes are possible. The definition of zero magnitude is arbitrary and corresponds to what was detectable when the first magnitude definition (C. F. Richter, 1935) was introduced.



The Local Magnitude (Richter Scale)

- Introduced by C. F. Richter in 1935.
- Symbol: M_L or ML
- X is the maximum amplitude A of a specific device, the Wood-Anderson seismometer.





The Wood-Anderson Seismometer

- Oscillation by torsion of a wire
- Electromagnetic damping
- Natural period of \approx 0.8 s (frequency $f_0=1.25\,\mathrm{Hz}$); close to the natural period of many building structures.



Relevant for earthquake hazard.

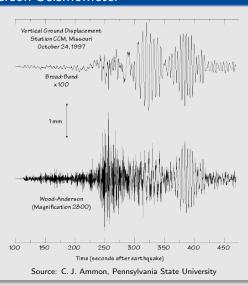
• Maximum magnification (record vs. ground displacement) of ≈ 2080 at f_0 ; sometimes a wrong value of 2800 was assumed.



Local magnitudes derived from synthesized seismograms were too high for some time.



The Wood-Anderson Seismometer





The Local Magnitude (Richter Scale)

• The local magnitude was originally defined as

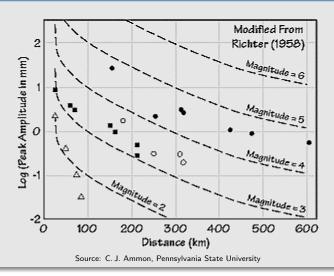
$$M_L = \log_{10} A \tag{61}$$

where the maximum amplitude A of the Wood-Anderson seismometer is measured in μ m at 100 km distance from the epicenter.

- $e = 1 \rightarrow 1$ unit increase in magnitude corresponds to an increase in the instrument's amplitude by a factor 10.
- Originally only a distance correction $\sigma(\Delta) = -\log_{10} A_0(\Delta)$ for shallow earthquakes ($h \le 15 \, \mathrm{km}$) in California was provided.

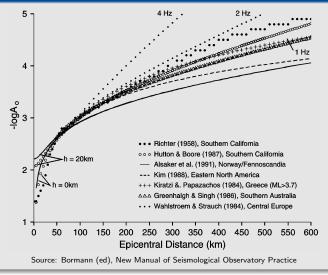


Richter's Original Distance Correction





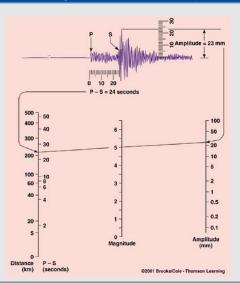
Distance Corrections for Different Regions and Depths



Intensity and Magnitude



Determining the Local Magnitude of an Earthquake





The Surface-Wave Magnitude

- Symbol: M_S or MS
- Original definition by B. Gutenberg (1945):

$$M_S = \log_{10} u_{h \max} + \sigma(\Delta) \tag{62}$$

where $u_{h\,\text{max}}$ is the maximum horizontal ground displacement at periods from $T=18\,\text{s}$ to $22\,\text{s}$.

Widely used modified definition (Moscow-Prague formula, 1962):

$$M_S = \max \left\{ \log_{10} \frac{|\vec{u}|}{T} \right\} + 1.66 \log_{10} \Delta + 3.3$$
 (63)

for $2^{\circ} \leq \Delta \leq 160^{\circ}.$ The maximum is taken over all periods of surface waves.



Body-Wave Magnitudes

- Two significantly different definitions
- Symbols: m_B , mB, m_b , mb,
- Original definition by B. Gutenberg (1945):

$$m_B = \max \left\{ \log_{10} \frac{|\vec{u}|}{T} \right\} + \sigma(\Delta)$$
 (64)

where $|\vec{u}|$ is analyzed for different types of body waves separately (with different functions $\sigma(\Delta)$ at periods from $T = 0.5 \,\mathrm{s}$ to 12 s.

• Alternative definition (m_b, mb) refers to higher-frequency components of P-waves only.



The Moment Magnitude

$$M_W = \frac{2}{3} \log_{10} M_0 - 6.1 \tag{65}$$

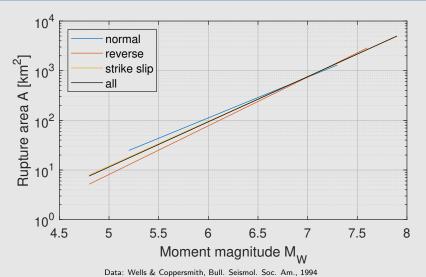
with M_0 in Nm

- Introduced in 1977 by H. Kanamori in order to characterize large earthquakes.
- More closely related to the strength of earthquakes at the seismic focus than older magnitude definitions.
- Rather a tectonic than a seismological magnitude scale.

Why is $e = \frac{2}{3}$ here while e = 1 for other magnitude scales?



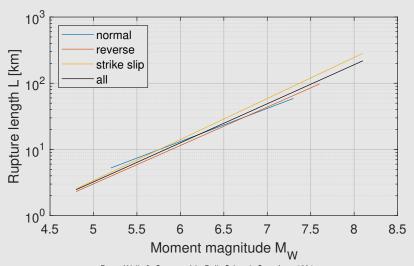
Scaling Properties of Earthquakes



80 /



Scaling Properties of Earthquakes



Data: Wells & Coppersmith, Bull. Seismol. Soc. Am., 1994



Scaling Properties of Earthquakes

Assume that the forces F in M_0 were uniformly distributed over an area A.

Shear stress

$$\sigma_s = \frac{F}{A} \tag{66}$$



Shear strain

$$\gamma = \frac{\sigma_s}{\mu} = \frac{F}{\mu A}$$



Displacement in shear direction

$$u = \gamma a = \frac{Fa}{\mu A} = \frac{M_0}{\mu A}$$

(68)



Scaling Properties of Earthquakes

Alternative interpretation of the seismic moment:

$$M_0 = \mu A \overline{u} \tag{69}$$

where

 $A = \text{size of the rupture area } [m^2]$

 \overline{u} = mean displacement along the rupture area [m]

Relation between A und \overline{u} ?



Scaling Properties of Earthquakes

Assume an elastic solid in a domain with a characteristic length scale L. Increase L, but keep displacement \vec{u} .

$$\epsilon = \frac{1}{2} \left(\nabla \vec{u} + (\nabla \vec{u})^T \right) \sim \frac{1}{L}$$
 (70)

$$\sigma = \lambda \, \epsilon_{\nu} \, \mathbf{1} + 2\mu \, \epsilon \, \sim \, \frac{1}{I} \tag{71}$$



Scaling Properties of Earthquakes

 σ remains constant if \vec{u} is also upscaled in the way $\vec{u} \sim L$.



$$\bar{u} \sim \sqrt{A}$$

(72)

if

- the stress drop (the component resulting from the forces in M_0) of all earthquakes is the same and
- the rupture areas of all earthquakes are geometrically similar.



$$\overline{u} \sim M_0^{\frac{1}{3}}, \quad A \sim M_0^{\frac{2}{3}}$$

(73)



(74)

(75)

Scaling Properties of Earthquakes

Rupture propagates along the rupture area at a given velocity



Duration
$$au \sim \sqrt{A} \sim M_0^{\frac{1}{3}}$$



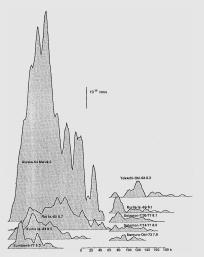
$$|\vec{u}| \sim \dot{M}_0 \sim \frac{M_0}{\tau} \sim M_0^{\frac{2}{3}}$$



Magnitude definition based on M_0 requires $e = \frac{2}{3}$.



Scaling Properties of Earthquakes



Source: Bormann (ed), New Manual of Seismological Observatory Practice

(76)

The Energy Magnitude

A crude scaling relation: particle velocity

$$|\vec{v}| \sim \frac{|\vec{u}|}{\tau} \sim \frac{M_0^{\frac{2}{3}}}{\tau} \sim M_0^{\frac{1}{3}}$$

Total radiated kinetic energy

Potential energy equals kinetic energy in the mean.

 $E_{\rm kin} \sim |\vec{v}|^2 \tau \sim M_0$



Total radiated seismic energy



 $E \sim M_0$



The Energy Magnitude

Theoretical relationship suggested by H. Kanamori (1977):

$$E \approx 5 \times 10^{-5} M_0 \tag{79}$$

Corresponding definition of the energy magnitude:

$$M_E = \frac{2}{3} \log_{10} \frac{E}{5 \times 10^{-5}} - 6.1 = \frac{2}{3} \log_{10} E - 3.2$$
 (80)

Up to one order of magnitude deviation from Kanamori's relationship was found for individual earthquakes.

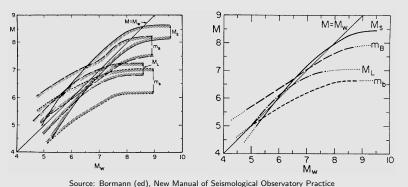


Significant differences between M_E and M_W for individual earthquakes.



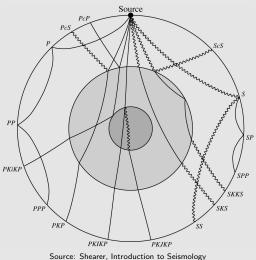
Saturation of Magnitudes

- All magnitudes based on recording seismic waves fall below M_W for large earthquakes.
- Effect is stronger if short-term (high-frequency) components of the seismic waves are used.



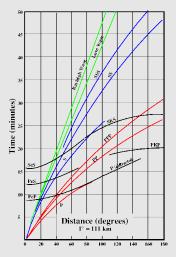


Global Wave Propagation in the Earth's Interior





Travel Time Curves



Source: Southern Arizona Seismic Observatory

Ray and Wave Front Approaches

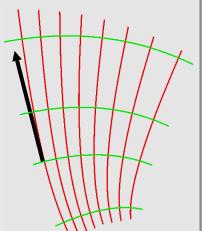
Plane wave approach and spherical wave solutions are only valid in a homogeneous medium.

Generalization:

Wave fronts are no longer planes, but arbitrary surfaces.

Direction of propagation is normal to the wave fronts

Ray paths are lines following the local direction of propagation (no longer straight lines).



Ray and Wave Front Approaches

Extensions towards the plane wave approach:

Replace

$$\vec{u}(\vec{x},t) = f(t - \vec{s} \cdot \vec{x}) \vec{a} \tag{81}$$

by

$$\vec{u}(\vec{x},t) = f(t - \psi(\vec{x})) \vec{s}(\vec{x})$$
 (82)

- Retarded time $\tau = t \psi(\vec{x})$ instead of $\tau = t \vec{s} \cdot \vec{x}$ with a general phase function $\psi(\vec{x})$
- Spatially variable amplitude vector $\vec{a}(\vec{x})$

Ray and Wave Front Approaches

V

Wave fronts: surfaces where $\psi(\vec{x})$ is constant

Direction of propagation: Wave propagates locally in direction of $\nabla \psi(\vec{x})$.

Slowness: $\vec{s}(\vec{x}) = \nabla \psi(\vec{x})$ defines the local slowness vector.

Ray paths: lines following the direction of $\vec{s} = \nabla \psi$

Limitation: Only valid in the limit of high frequencies.

Ray and Wave Front Approaches



Main results:

Types of waves: two independent types of waves (P-wave and S-wave); amplitude vector \vec{a} is either parallel or normal to $\vec{s} = \nabla \psi$ at each point; no reflection; no merging of both wave types

Slowness/velocity: same relationship as for plane and spherical waves (eikonal equation)

$$|\vec{s}|^2 = |\nabla \psi|^2 = \frac{\rho}{\lambda + 2\mu}$$
 or $|\vec{s}|^2 = |\nabla \psi|^2 = \frac{\rho}{\mu}$ (83)

Ray and Wave Front Approaches

V.

Amplitudes: Absolute values of the amplitudes are determined by the condition

$$\operatorname{div}(\vec{q}) = 0 \tag{84}$$

with the flux density

$$\vec{q} = \rho |\vec{a}|^2 \vec{v} = \epsilon \vec{v} \tag{85}$$

and $\epsilon = \rho |\vec{a}|^2$.

 ϵ is proportional to the energy density (elastic energy + kinetic energy per volume) of a harmonic wave.



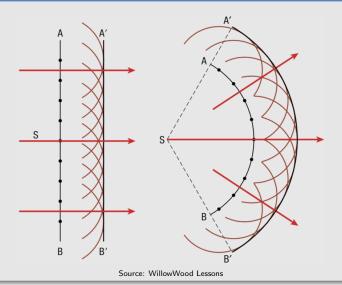
Call ϵ energy density and \vec{q} energy flux density.

Geometrical and Numerical Approaches

- Numerical solution of the eikonal equation: one-sided difference quotients for $\psi(\vec{x})$ from the considered points to neighbors where ψ is already known; similar to upstream scheme.
- Huygens' princple: Construct wave fronts (planes where $\psi(\vec{x})$ is constant) progressively for increasing values of ψ ; rather a geometrical approach than a numerical method.
- Ray tracing: Compute individual rays (lines in dierection of $\nabla \psi(\vec{x})$).



Construction of Wave Fronts – Huygens' Princple



Construction of Ray Paths

Each ray path $\vec{r}(t)$ must be normal to the wave fronts ($\phi = \text{const}$).



$$\frac{d}{dt}\vec{r}(t) \sim \nabla \psi(\vec{r}(t)) = \vec{s}(\vec{r}(t))$$
 (86)

Follow the slowness vector along the ray path:

$$\frac{d}{dt}\vec{s}(\vec{r}(t)) = \nabla \vec{s}(\vec{r}(t)) \frac{d}{dt}\vec{r}(t) \tag{87}$$

$$\sim \nabla \vec{s}(\vec{r}(t)) \vec{s}(\vec{r}(t)) \tag{88}$$

$$= \frac{1}{2} \nabla |\vec{s}(\vec{r}(t))|^2 \text{ because } \nabla \vec{s}^T = \nabla \vec{s} \tag{89}$$

$$= |\vec{s}(\vec{r}(t))| \nabla |\vec{s}(\vec{r}(t))| \tag{90}$$

Construction of Ray Paths

 \vec{s} changes along the ray path in direction of increasing slowness (decreasing velocity).

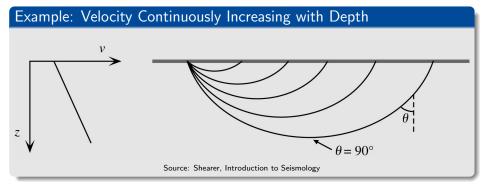


Component of \vec{s} parallel to the surfaces of constant slowness remains constant.



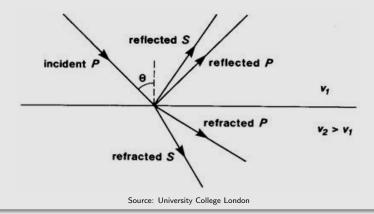
Horizontal slowness remains constant.

Horizontal direction in seismology = planes where the properties of the material are constant.



Reflection and Refraction

Simplest case: two homogeneous, isotropic halfspaces with different properties (λ, μ, ρ) and plane waves in each of them; horizontal interface at $x_3 = 0$.



Reflection and Refraction



Each wave is considered as a plane wave

$$\vec{u}(\vec{x},t) = f(t - \vec{s} \cdot \vec{x}) \vec{a} \tag{91}$$

with individual values \vec{s} and \vec{a} where either

$$\vec{a} \parallel \vec{s}$$
 and $|\vec{s}| = s_p = \sqrt{\frac{\rho}{\lambda + 2\mu}}$ (92)

or

$$\vec{a} \perp \vec{s}$$
 and $|\vec{s}| = s_s = \sqrt{\frac{\rho}{\mu}}$ (93)

Reflection and Refraction

Displacement must be continuous at $x_3 = 0$.



All waves must have the same retarded time

$$\tau = t - \vec{s} \cdot \vec{x} = t - s_1 x_1 - s_2 x_2 \tag{94}$$

at the interface.



Components of \vec{s} parallel to the interface (s_1, s_2) must be the same for all waves.



Horizontal slowness remains constant in reflection and refraction.

Reflection and Refraction



Second condition at the interface:

Stress acting on the interface, given by

$$\vec{\sigma}_{\text{int}} = \sigma|_{x_3=0} \begin{pmatrix} 0\\0\\1 \end{pmatrix} \tag{95}$$

must be continuous.

Continuity of displacement and stress at the interface



Linear equation system for the amplitude vectors \vec{a} of the five involved waves.

Conversion of Waves in Reflection and Refraction

<u>V</u>

Align the coordinate system in such a way that all waves propagate in the x_1 - x_3 plane ($s_2 = 0$, possible because s_1 and s_2 are the same for all involved waves).

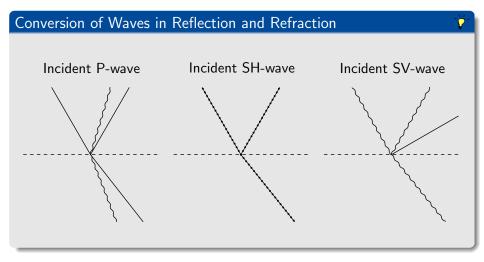


Vertically polarized S-wave (SV-wave):

- $a_2 = 0$ \Longrightarrow particle displacement in the x_1 - x_3 plane (and normal to wave propagation)
- Converted to (and from) P and SV waves in reflection and refraction

Horizontally polarized S-wave (SH-wave):

- $a_1 = a_3 = 0$ \longrightarrow particle displacement in x_2 direction
- Independent of P and SV waves





The Waves Disappearing in Reflection and Refraction



For all waves involved in refection and refraction at a planar interface $(s_2 = 0)$: s_1 given and

$$|\vec{s}|^2 = s_1^2 + s_3^2 = \begin{cases} \frac{1}{v_\rho^2} = \frac{\rho}{\lambda + 2\mu} & \text{P-waves} \\ \frac{1}{v_s^2} = \frac{\rho}{\mu} & \text{for S-waves} \end{cases}$$
 (96)

given.



Waves cannot propagate if $s_1 > |\vec{s}|$, but

$$s_3 = \pm \sqrt{|\vec{s}|^2 - s_1^2} = \pm i \sqrt{s_1^2 - |\vec{s}|^2}$$
 (97)

would be a formal solution.



Harmonic Interface Waves



For a harmonic wave:

$$\vec{u}(\vec{x},t) = e^{i\omega(t-\vec{s}\cdot\vec{x})}\vec{a} = e^{i\omega(t-s_1x_1+i\sqrt{s_1^2-|\vec{s}|^2}x_3)}\vec{a}$$

$$= e^{i\omega(t-s_1x_1)}e^{\pm\omega\sqrt{s_1^2-|\vec{s}|^2}x_3}\vec{a}$$
(98)
(99)

Can be considered as a wave propagating along the interface (here in x_1 direction) with an amplitude depending on x_3 :

$$\vec{u}(\vec{x},t) = e^{i\omega(t-s_1x_1)}\vec{a}_{eff}$$
 (100)

with the effective amplitude

$$\vec{a}_{\text{eff}} = e^{\pm\omega\sqrt{s_1^2 - |\vec{s}|^2} x_3} \vec{a} \tag{101}$$



Harmonic Interface Waves



Only the version where \vec{a}_{eff} decreases exponentially with distance from the interface ($|x_3|$) makes sense:

$$\vec{a}_{\text{eff}} = e^{-\omega\sqrt{s_1^2 - |\vec{s}|^2}|x_3|} \vec{a}$$
 (102)

Respective waves are called interface waves.



The Depth of Penetration of Interface Waves



$$\vec{a}_{\text{eff}} = e^{-\omega \sqrt{s_1^2 - |\vec{s}|^2} |x_3|} \vec{a} = e^{-\frac{|x_3|}{d}} \vec{a}$$
 (103)

with the depth of penetration

$$d = \frac{1}{\omega \sqrt{s_1^2 - |\vec{s}|^2}} = \frac{L}{2\pi \sqrt{1 - \frac{|\vec{s}|^2}{s_1^2}}}$$
(104)

and the wavelength $L = \frac{2\pi}{\omega s_1}$.

- $d \to \infty$ (plane wave propagating along the surface) if the wave is only slightly too slow for the medium $(s_1 \to |\vec{s}|)$.
- $d o \frac{L}{2\pi}$ if the wave is much too slow for the medium $(s_1 \gg |\vec{s}|)$.



Particle Orbits of P and SV Interface Waves



Examples of particle orbits for an incident SV wave at the crust-mantle boundary: $\alpha=20^\circ$, $\alpha=30^\circ$, $\alpha=40^\circ$, $\alpha=70^\circ$

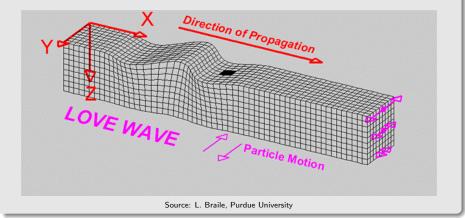
General properties:

- Particles move on elliptical orbits.
- Horizontal ellipses for P interface waves, vertical ellipses for S interface waves.
- Aspect ratio of the ellipses $= \frac{1}{\sqrt{1-\frac{|\vec{s}|^2}{s_1^2}}}.$
- Prograde rotation in the lower halfspace; retrograde rotation in the upper halfspace.

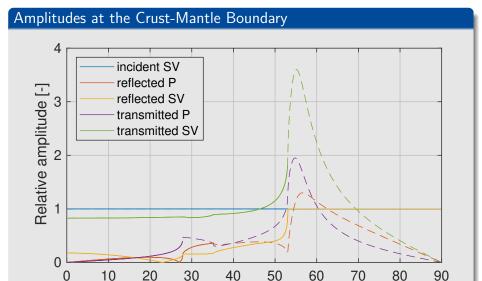


Particle Orbits of SH Interface Waves

Particles move on straight lines parallel to the interface (same as for regular SH waves).







 α [°]



Surface Waves at a Free Surface

Interface waves: driven by plane waves (incident, reflected, refracted)

Surface waves: living on their own at a free surface (interface to air);

$$\vec{\sigma}_{\mathsf{surf}} = \vec{\mathsf{0}}$$



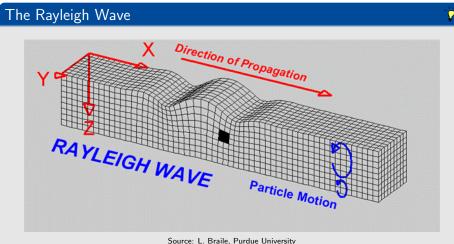
Two fundamental types of surface waves in a semi-infinite halfspace:

Rayleigh wave, named after J. W. Strutt (later 3. Lord Rayleigh)

Love wave, named after A. E. H. Love

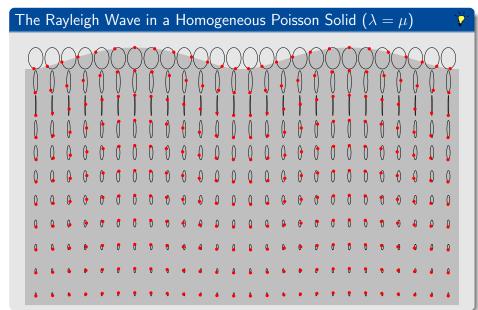
Not possible in a homogeneous halfspace because $\vec{\sigma}_{\mathsf{surf}} \neq \vec{0}$ for $a_2 \neq 0$





Specific superposition of P and SH interface wave so that $\vec{\sigma}_{\text{surf}} = \vec{0}$







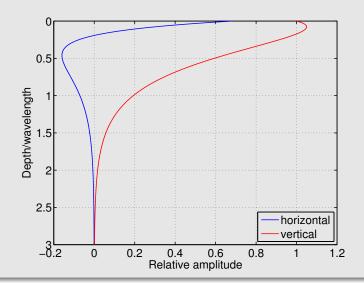
The Rayleigh Wave in a Homogeneous Poisson Solid $(\lambda = \mu)$



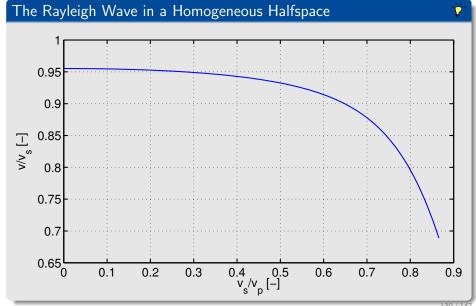
- Retrograde particle motion on elliptical orbits at the surface.
- Prograde particle motion on elliptical orbits at greater depth.
- Velocity $v \approx 0.92 v_s$.



The Rayleigh Wave in a Homogeneous Poisson Solid $(\lambda = \mu)$









(105)

Surface Waves in Inhomogeneous Media

Assume that ρ , λ , and μ depend on the vertical coordinate x_3 , and generalize

reralize
$$\vec{u}(\vec{x},t) = e^{i\omega(t-s_1x_1)} \left(e^{k_1S_px_3} a_p \begin{pmatrix} 1\\0\\iS_p \end{pmatrix} + e^{k_1S_sx_3} a_s \begin{pmatrix} 1\\0\\\frac{i}{r} \end{pmatrix} \right)$$

to

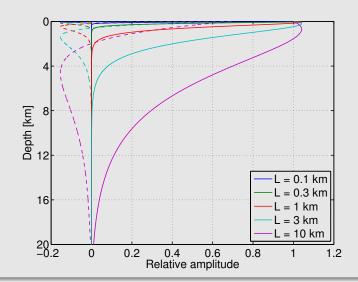
$$\vec{u}(\vec{x},t) = e^{i\omega(t-s_1x_1)}\vec{a}(x_3)$$
 (106)



- Differential equations (eigenvalue problem) for $a_1(x_3)$, $a_2(x_3)$, $a_3(x_3)$.
- $a_1(x_3)$ and $a_3(x_3)$ are coupled (\rightarrow Rayleigh wave), $a_2(x_3)$ is independent of $a_1(x_3)$ and $a_3(x_3)$ (\rightarrow Love wave).
- Must be solved numerically.

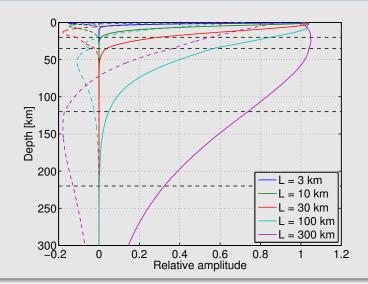


Rayleigh Waves in Typical Continental Subsurface (PEM)



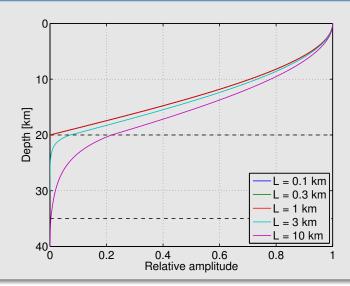


Rayleigh Waves in Typical Continental Subsurface (PEM)



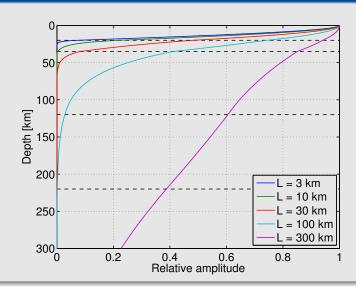


Love Waves in Typical Continental Subsurface (PEM)





Love Waves in Typical Continental Subsurface (PEM)





Main Differences Between Body Waves and Surface Waves

• Decrease with the distance from the hypocenter/epicenter r:

	Energy flux density	Amplitude
body waves	$\propto \frac{1}{r^2}$	$\propto \frac{1}{r}$
surface waves	$\propto \frac{1}{r}$	$\propto \frac{1}{\sqrt{r}}$



Surface waves have a longer range than body waves.

 Velocity of harmonic surface waves depends on the wavelength (dispersion).

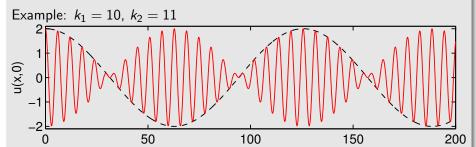


Dispersion

Simplest situation: superposition of two harmonic waves with the same amplitude (=1), but different frequencies in 1D:

$$u(x,t) = e^{i\omega_1(t-s_1x)} + e^{i\omega_2(t-s_2x)} = e^{i(\omega_1t-k_1x)} + e^{i(\omega_2t-k_2x)}$$
 (107)

with $k_1 = \omega_1 s_1$, $k_2 = \omega_2 s_2$.





Dispersion

$$u(x,t) = e^{i\omega_{1}(t-s_{1}x)} + e^{i\omega_{2}(t-s_{2}x)} = e^{i(\omega_{1}t-k_{1}x)} + e^{i(\omega_{2}t-k_{2}x)}$$
(108)

$$= e^{i\left(\overline{\omega}t + \frac{\omega_{1}-\omega_{2}}{2}t - \overline{k}x - \frac{k_{1}-k_{2}}{2}x\right)} + e^{i\left(\overline{\omega}t - \frac{\omega_{1}-\omega_{2}}{2}t - \overline{k}x + \frac{k_{1}-k_{2}}{2}x\right)}$$
(109)

$$= e^{i\left(\overline{\omega}t - \overline{k}x\right)} \left(e^{i\left(\frac{\omega_{1}-\omega_{2}}{2}t - \frac{k_{1}-k_{2}}{2}x\right)} + e^{i\left(-\frac{\omega_{1}-\omega_{2}}{2}t + \frac{k_{1}-k_{2}}{2}x\right)}\right)$$
(110)

$$= e^{i\left(\overline{\omega}t - \overline{k}x\right)} 2\cos\left(\frac{\omega_{1}-\omega_{2}}{2}t - \frac{k_{1}-k_{2}}{2}x\right)$$
(111)

$$= e^{i\overline{\omega}\left(t - \frac{\overline{k}}{\overline{\omega}}x\right)} 2\cos\left(\frac{\omega_{1}-\omega_{2}}{2}\left(t - \frac{k_{1}-k_{2}}{\omega_{1}-\omega_{2}}x\right)\right)$$
(112)

with $\overline{\omega} = \frac{\omega_1 + \omega_2}{2}$, $\overline{k} = \frac{k_1 + k_2}{2}$.



Dispersion

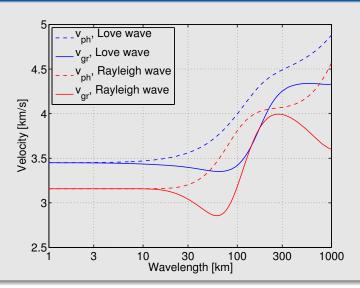
For $\omega_1 \approx \omega_2$, $k_1 \approx k_2$:

- High-frequency oscillation with an angular frequency $\overline{\omega}$ propagating with the phase slowness $s_{\rm ph}=\frac{\overline{k}}{\overline{\omega}}$
- Low-frequency oscillation of the amplitude with an angular frequency $\frac{\omega_1-\omega_2}{2}$ propagating with the group slowness

$$s_{\rm gr} = \frac{k_1 - k_2}{\omega_1 - \omega_2} \rightarrow \frac{dk}{d\omega} \quad \text{for} \quad \omega_1 - \omega_2 \rightarrow 0$$
 (113)

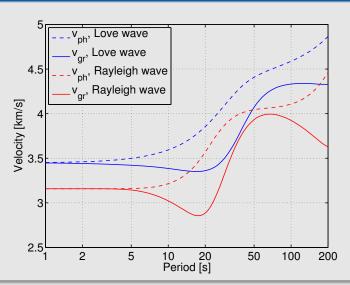


Velocities of Surface Waves in Typical Continental Subsurface (PEM)

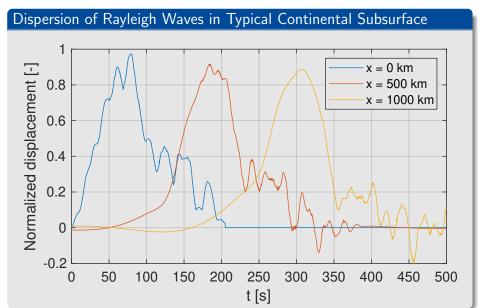




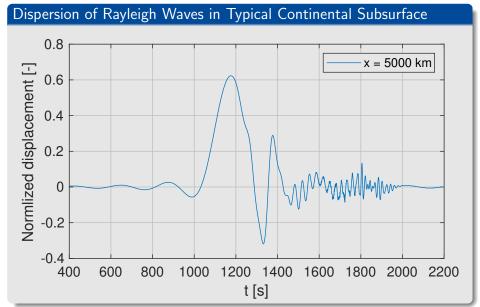
Velocities of Surface Waves in Typical Continental Subsurface (PEM)



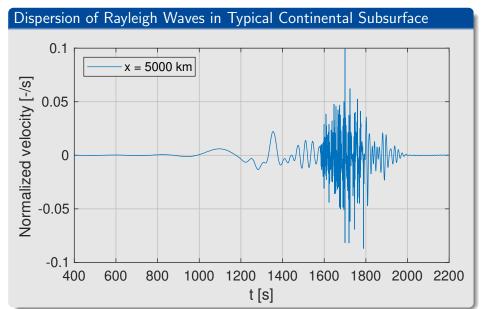




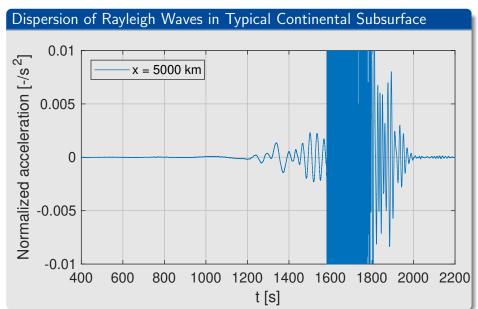






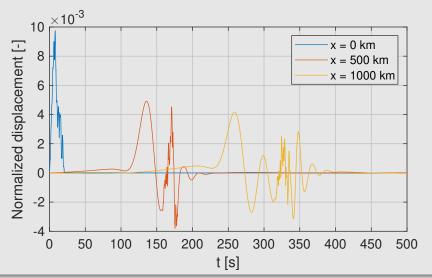








Dispersion of Rayleigh Waves in Typical Continental Subsurface





Dispersion of Rayleigh Waves in Typical Continental Subsurface

