

Tsunamis

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Main Properties of Tsunamis

- Gravity waves in oceans with long periods between about 100 s and 10,000 s.
- Tsunamis propagate at high velocities in deep water.
- Mainly horizontal particle motion involving the entire water column down to the ocean floor.



Rather small dissipation of energy.



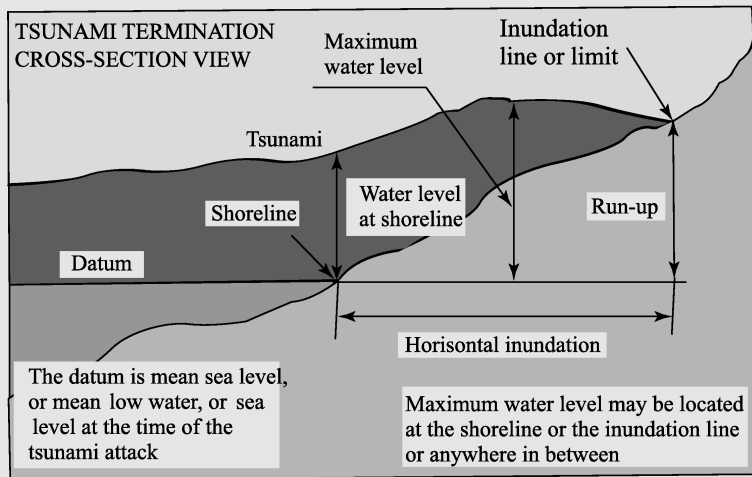
Tsunamis travel over large distances.

- Wave height increases with decreasing ocean depth.



Tsunamis may reach large wave heights at the coast.

Basic Terms

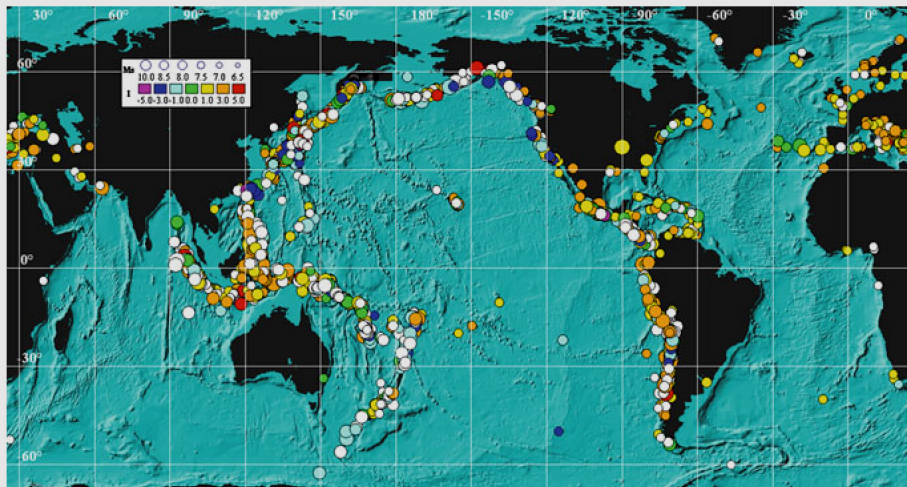


Source: Levin & Nosov, Physics of Tsunamis

Main Sources of Tsunamis

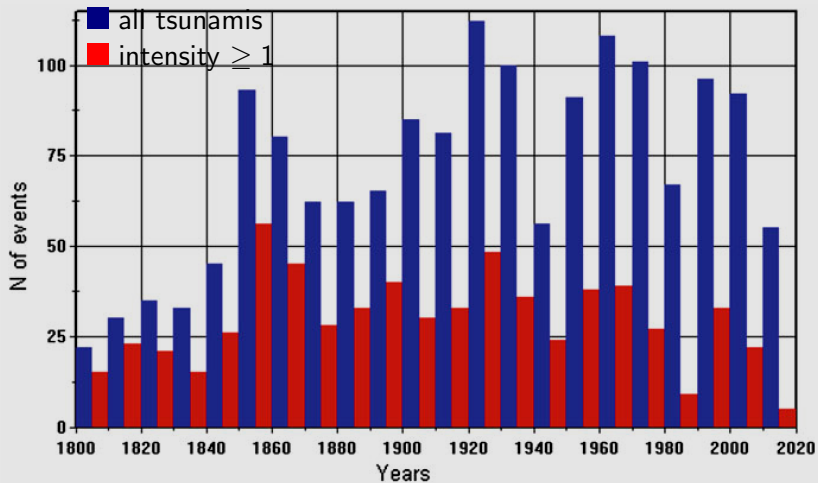
- Earthquakes (more than 90 % of all tsunamis)
- Landslides
- Volcanic eruptions
- Meteorite impact (rare)

Worldwide Distribution of Tsunami Sources from 2000 B.C. to 2014



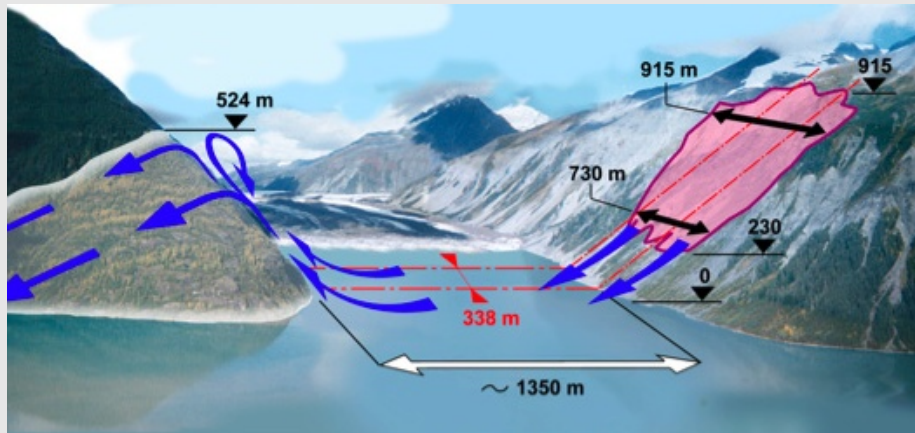
Source: Levin & Nosov, Physics of Tsunamis

Worldwide Number of Tsunamis per Decade



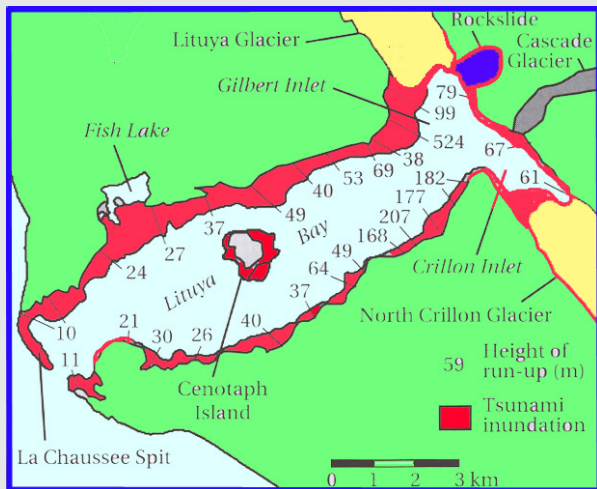
Source: Levin & Nosov, Physics of Tsunamis

The Tallest Tsunami Known so far: Lituya Bay, 1958



Source: Pararas-Carayannis, The Mega-Tsunami of July 9, 1958 in Lituya Bay, Alaska

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The Tallest Tsunamis 2000–2014

Date	Location	M_W	H_{\max} [m]	Death toll
11.03.2011	Japan	9.0	56	18,482
24.12.2004	Indonesia, Sumatra	9.1	51	227,899
27.02.2010	Chile	8.8	29	156
29.09.2009	Samoa	8.1	22	192
15.11.2006	Russia, Kuril Islands	8.3	22	0
17.07.2006	Indonesia, South of Java	7.7	21	802
25.10.2010	Indonesia, Sumatra	7.8	17	431

Types of Intensity and Magnitude Scales

Three different types of scales:

- Intensity scales characterizing the effect of a tsunami on humans and their structures (Sieberg-Ambraseys scale, Papadopoulos-Imamura scale).
- Intensity scales based on measurements of wave height at the coast (Imamura-Iida scale, Soloviev-Imamura scale).
- Magnitude scales characterizing the strength of a tsunami independent of distance between source and coast and the shape of the coast (Abe-Hatori scale, Murty-Loomis scale).

The Sieberg-Ambraseys Scale

- Originally introduced by A. H. Sieberg (1927), modified by N. N. Ambraseys (1962).
- Six-point scale from 1 = very light to 6 = disastrous.

The Papadopoulos-Imamura Scale

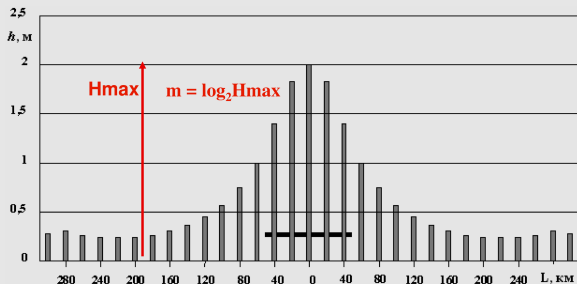
- Introduced by G. A. Papadopoulos and F. Imamura (2001).
- 12-point scale in analogy to the Mercalli scale for earthquakes from I = not felt to XII = destructive.

The Imamura-Iida Scale

- Introduced by A. Imamura (1942), modified by K. Iida (1956).
- Defined as

$$m = \log_2 H_{\max} \quad (1)$$

where H_{\max} is the maximum wave height.



Source: Gusiakov, Tsunami Quantification

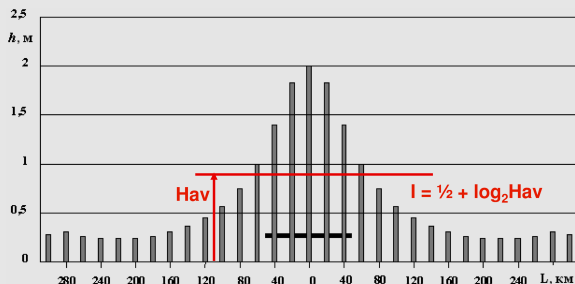
- Originally termed magnitude.

The Soloviev-Imamura Scale

- Modification of the Imamura-Iida scale by S. Soloviev (1972).
- Defined as

$$I = \frac{1}{2} + \log_2 H_{av} \quad (2)$$

where H_{av} is the average wave height along the nearest coast.



Source: Gusiakov, Tsunami Quantification

- Widely used in many tsunami catalogs.

The Abe-Hatori Scale

- Introduced in 1979 by K. Abe.
- First attempt to define a tsunami magnitude taking into account the distance from the source:

$$M_t = a \log_{10} H_{\max} + b \log_{10} \Delta + D \quad (3)$$

where

H_{\max} = maximum wave amplitude at the coast

Δ = distance

a, b, D = constants

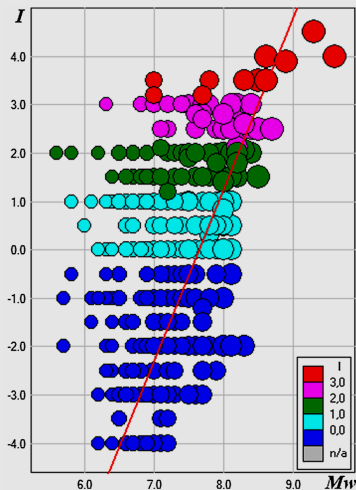
The Murty-Loomis Scale

- Introduced in 1980 by T. S. Murty and H. G. Loomis.
- Based on the total potential energy E (in J here, originally in ergs):

$$ML = 2(\log_{10} E - 12) \quad (4)$$

- Well- defined and
- theoretically a good measure of the strength of a tsunami,
- but suffers from the problem of determining the total potential energy.

Tsunami Intensity (Soloviev-Imamura) vs. Earthquake Magnitude



Source: Gusiakov, Pure Appl. Geophys, 2015

Starting Point

Elastic medium with $\mu = 0$:

$$\boldsymbol{\sigma} = \lambda \epsilon_v \mathbf{1} = -p \mathbf{1} \quad (5)$$

where

$$p = -\lambda \epsilon_v = -\lambda \operatorname{div}(\vec{u}) \quad (6)$$

is called **pressure**.

- Mechanically equivalent to a compressible, inviscid fluid.
- Theory is only valid for small displacement.



Kinematically not appropriate for describing fluids in general, but for waves with small amplitudes.

Types of Waves in Fluids

P body waves: sound wave; slowness

$$s = \frac{1}{v_p} = \sqrt{\frac{\rho}{\lambda}} \quad (7)$$

P interface waves:

- Prograde particle rotation on horizontal elliptical orbits in the lower halfspace.
- Amplitude decreases exponentially with depth; depth of penetration

$$d = \frac{1}{\omega \sqrt{s_1^2 - |\vec{s}|^2}} = \frac{L}{2\pi \sqrt{1 - \frac{|\vec{s}|^2}{s_1^2}}} \quad (8)$$

with the wavelength $L = \frac{2\pi}{\omega s_1}$.

P Interface Waves in Fluids



Particle displacement:

$$\vec{u}(\vec{x}, t) = e^{i\omega(t - \vec{s} \cdot \vec{x})} \vec{a} \quad (9)$$

$$= e^{i\omega(t - s_1 x_1)} e^{\pm \frac{x_3}{d}} \vec{a} \quad (10)$$

with

$$\frac{u_3(\vec{x}, t)}{u_1(\vec{x}, t)} = \frac{a_3}{a_1} = \frac{s_3}{s_1} \quad (11)$$

$$= \frac{\pm i \sqrt{s_1^2 - |\vec{s}|^2}}{s_1} = \pm i \sqrt{1 - \frac{|\vec{s}|^2}{s_1^2}} = \pm \frac{i}{\omega s_1 d} \quad (12)$$

where only the + sign makes sense in the lower halfspace.

P Interface Waves in Fluids

Pressure:

$$p(\vec{x}, t) = -\lambda \operatorname{div}(\vec{u}(\vec{x}, t)) = i\omega \lambda \vec{s} \cdot \vec{a} e^{i\omega(t - \vec{s} \cdot \vec{x})} \quad (13)$$

From Eq. 12:

$$\begin{aligned} \vec{s} \cdot \vec{a} &= s_1 a_1 + s_3 a_3 = s_1 a_1 \left(1 + \frac{s_3 a_3}{s_1 a_1} \right) = s_1 a_1 \left(1 - \left(1 - \frac{|\vec{s}|^2}{s_1^2} \right) \right) \\ &= \frac{a_1 |\vec{s}|^2}{s_1} = a_1 \frac{\rho}{\lambda s_1} = \pm a_3 \frac{\rho \omega d}{i \lambda} \end{aligned} \quad (14)$$
$$(15)$$



$$p(\vec{x}, t) = \pm \rho \omega^2 d a_3 e^{i\omega(t - \vec{s} \cdot \vec{x})} = \pm \rho \omega^2 d u_3(\vec{x}, t) \quad (16)$$

P Surface Waves in Fluids

No surface with $p(\vec{x}, t) = 0$ (or constant) at any time



P interface wave cannot be a surface wave.

P Surface Waves in Fluids with Gravity

Gravity causes additional hydrostatic pressure

$$p_{\text{hy}}(\vec{x}, t) = -\rho g(x_3 + u_3(\vec{x}, t)) \quad (17)$$



$$p(\vec{x}, t) = \pm \rho \omega^2 d u_3(\vec{x}, t) - \rho g(x_3 + u_3(\vec{x}, t)) \quad (18)$$

P Surface Waves in Fluids with Gravity

Free surface at $x_3 = 0$ ($p(x_1, x_2, 0, t) = 0$) is possible if

$$\omega^2 = \frac{g}{d} = \omega g \sqrt{s_1^2 - |\vec{s}|^2} \quad (19)$$

The Velocity of Propagation

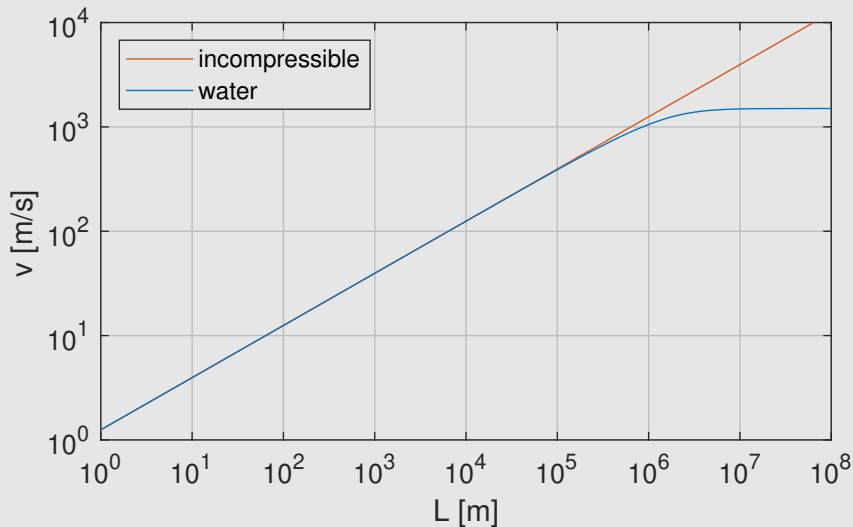
Slowness:

$$s_1^2 = |\vec{s}|^2 + \left(\frac{\omega}{g}\right)^2 \quad (20)$$

Often expressed in terms of the wavelength $L = \frac{2\pi}{\omega s_1}$:

$$s_1^2 = \frac{|\vec{s}|^2}{2} + \sqrt{\left(\frac{|\vec{s}|^2}{2}\right)^2 + \left(\frac{2\pi}{gL}\right)^2} \quad (21)$$

The Velocity of Propagation



Boundary Condition at the Ocean Floor

Consider domain $-H \leq x_3 \leq 0$ with a given ocean depth H .



Solution must meet the condition $u_3(x_1, x_2, -H, t) = 0$.

Superposition of the solutions with $+$ and $-$ signs with factors $\pm e^{\pm \frac{H}{d}}$

$$u_3(\vec{x}, t) = e^{\frac{H}{d}} \left(a_3 e^{i\omega(t-s_1x_1)} e^{\frac{x_3}{d}} \right) - e^{-\frac{H}{d}} \left(a_3 e^{i\omega(t-s_1x_1)} e^{-\frac{x_3}{d}} \right) \quad (22)$$

$$= a_3 e^{i\omega(t-s_1x_1)} \left(e^{\frac{x_3+H}{d}} - e^{-\frac{x_3+H}{d}} \right) \quad (23)$$

satisfies the boundary condition at $x_3 = -H$.

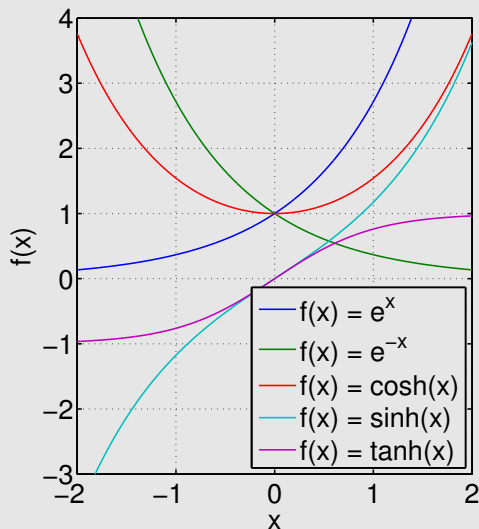
The Hyperbolic Cosine, Sine and Tangent Functions

$$\cosh(x) = \frac{e^x + e^{-x}}{2} \quad (24)$$

$$\sinh(x) = \frac{e^x - e^{-x}}{2} \quad (25)$$

$$\tanh(x) = \frac{\sinh(x)}{\cosh(x)} \quad (26)$$

$$= \frac{e^x - e^{-x}}{e^x + e^{-x}} \quad (27)$$



Vertical Particle Displacement

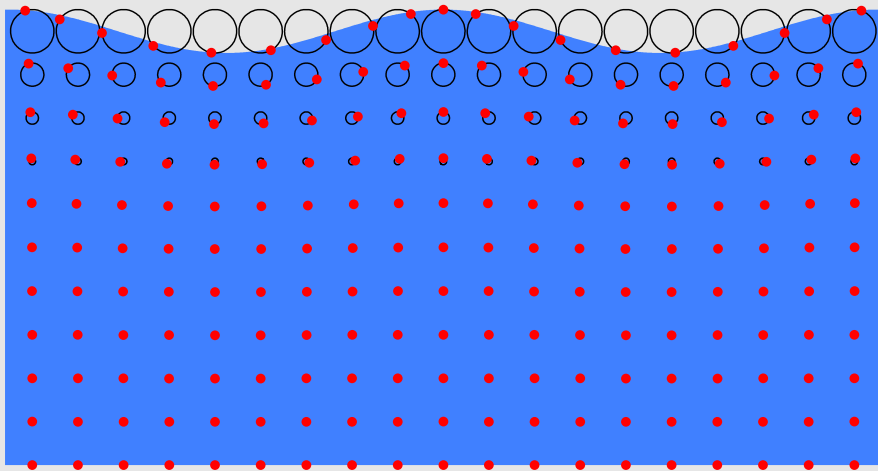
Wave height

$$h = u_3(\vec{0}, 0) = a_3 \left(e^{\frac{H}{d}} - e^{-\frac{H}{d}} \right) = 2a_3 \sinh \left(\frac{H}{d} \right) \quad (28)$$

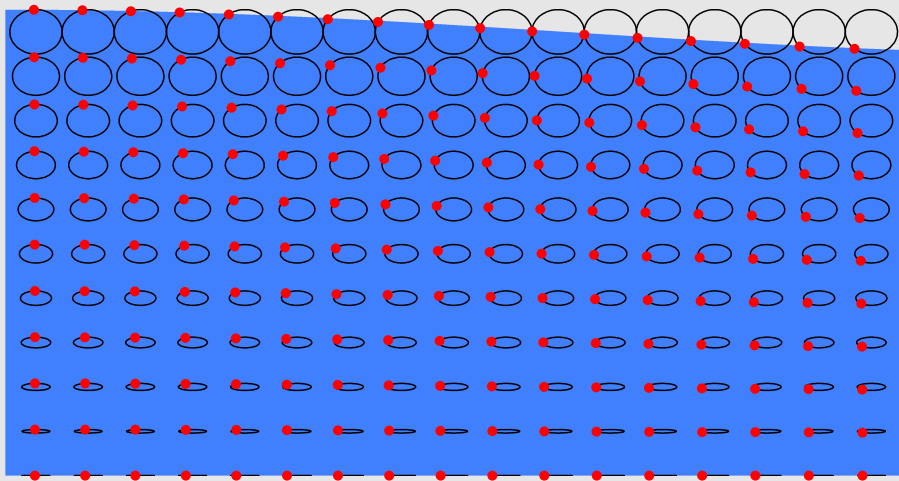


$$u_3(\vec{x}, t) = h e^{i\omega(t-s_1x_1)} \frac{\sinh \left(\frac{x_3+H}{d} \right)}{\sinh \left(\frac{H}{d} \right)} \quad (29)$$

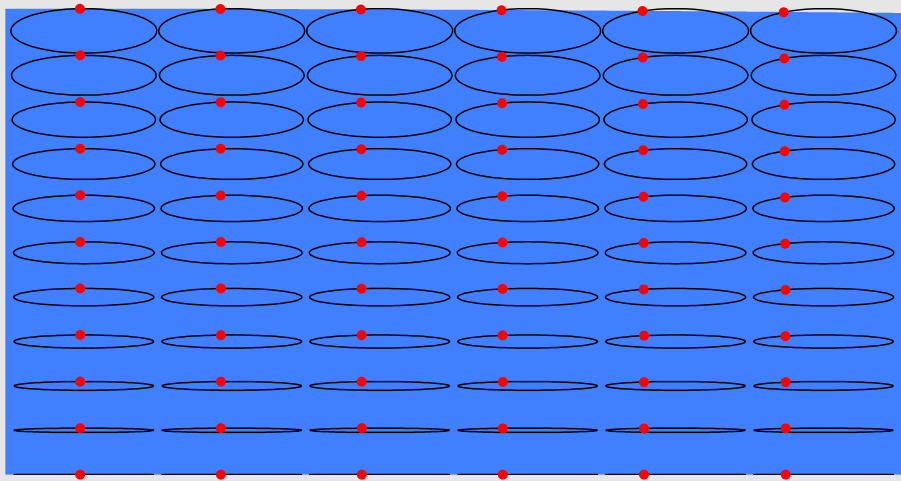
Particle Orbits for $L/H = 1$ (incompressible)



Particle Orbits for $L/H = 5$ (incompressible)



Particle Orbits for $L/H = 20$ (incompressible)



The Velocity of Propagation

Pressure (Eq. 18) for the superposed solution (Eq. 23):

$$\begin{aligned} p(\vec{x}, t) &= \rho\omega^2 d \left(a_3 e^{i\omega(t-s_1x_1)} \left(e^{\frac{x_3+H}{d}} + e^{-\frac{x_3+H}{d}} \right) \right) \\ &\quad - \rho g \left(x_3 + \left(a_3 e^{i\omega(t-s_1x_1)} \left(e^{\frac{x_3+H}{d}} - e^{-\frac{x_3+H}{d}} \right) \right) \right) \end{aligned} \quad (30)$$

$$\begin{aligned} &= 2a_3\rho e^{i\omega(t-s_1x_1)} \left(\omega^2 d \cosh\left(\frac{x_3+H}{d}\right) - g \sinh\left(\frac{x_3+H}{d}\right) \right) \\ &\quad - \rho g x_3 \end{aligned} \quad (31)$$

Free surface at $x_3 = 0$ ($p(x_1, x_2, 0, t) = 0$) is possible if

$$\omega^2 = \frac{g}{d} \tanh\left(\frac{H}{d}\right) = \omega g \sqrt{s_1^2 - |\vec{s}|^2} \tanh\left(\frac{H}{d}\right) \quad (32)$$

The Velocity of Propagation

Generalization of Eq. 21:

$$s_1^2 = \frac{|\vec{s}|^2}{2} + \sqrt{\left(\frac{|\vec{s}|^2}{2}\right)^2 + \left(\frac{2\pi}{gL \tanh\left(\frac{H}{d}\right)}\right)^2} \quad (33)$$

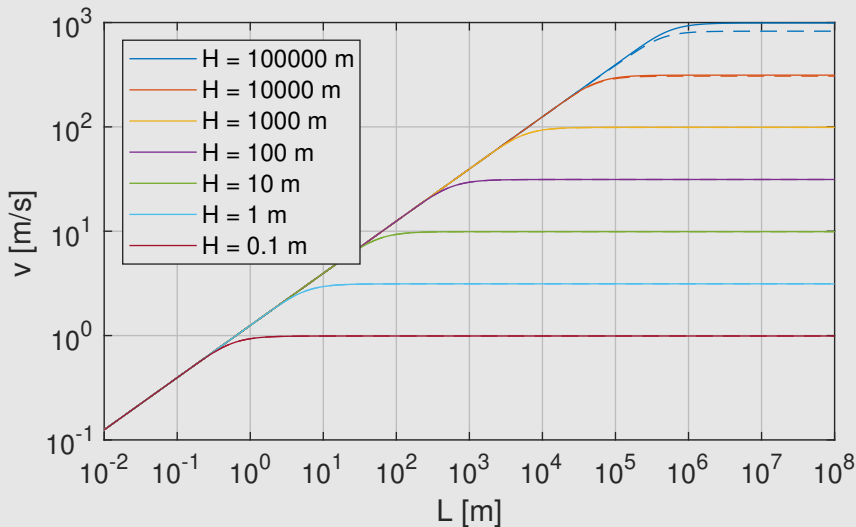
For incompressible fluids ($|\vec{s}| = 0$, $d = \frac{L}{2\pi}$):

$$s_1^2 = \frac{2\pi}{gL \tanh\left(\frac{2\pi H}{L}\right)} \quad (34)$$



$$v = \sqrt{\frac{gL}{2\pi} \tanh\left(\frac{2\pi H}{L}\right)} \quad (35)$$

The Velocity of Propagation



Regimes of Ocean Wave Propagation

Deep water regime: $\frac{d}{H} \leq \frac{1}{\pi} \Leftrightarrow \frac{L}{H} \leq 2$

- Particles move on almost circular orbits.
- Particle movement is practically limited to a depth less than one wavelength.
- Horizontal particle displacement at the ocean floor is less than 10% of the displacement at the surface.
- Velocity depends on the wavelength, but not on ocean depth:

$$v \approx \sqrt{\frac{gL}{2\pi}}, \quad (36)$$

- Strong dispersion

Regimes of Ocean Wave Propagation

Shallow water regime: $\frac{d}{H} \geq \frac{10}{\pi} \Leftrightarrow \frac{L}{H} \geq 20$

- Particles move on elliptical orbits.
- Horizontal particle movement persists down to the ocean floor; at the ocean floor more than 95 % of the displacement at the surface.
- Velocity only depends on ocean depth:

$$v \approx \sqrt{gH}$$

- No dispersion

Dispersion

Examples of tsunami wave dispersion in a 4000 m deep ocean (symmetric propagation to the left and to the right):

- bell-shaped (Gaussian) wave
- boxcar-shaped wave
- double boxcar-shaped wave
- step-like wave

The Fluid Pressure

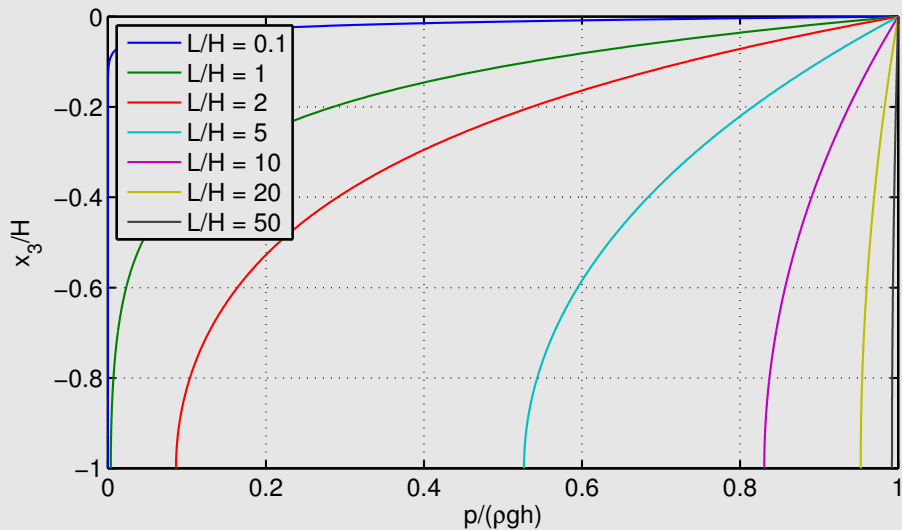
Variation in fluid pressure without hydrostatic pressure from Eqs. 31, 28, and 32:

$$p(\vec{x}, t) = 2a_3\rho e^{i\omega(t-s_1x_1)} \omega^2 d \cosh\left(\frac{x_3+H}{d}\right) \quad (37)$$

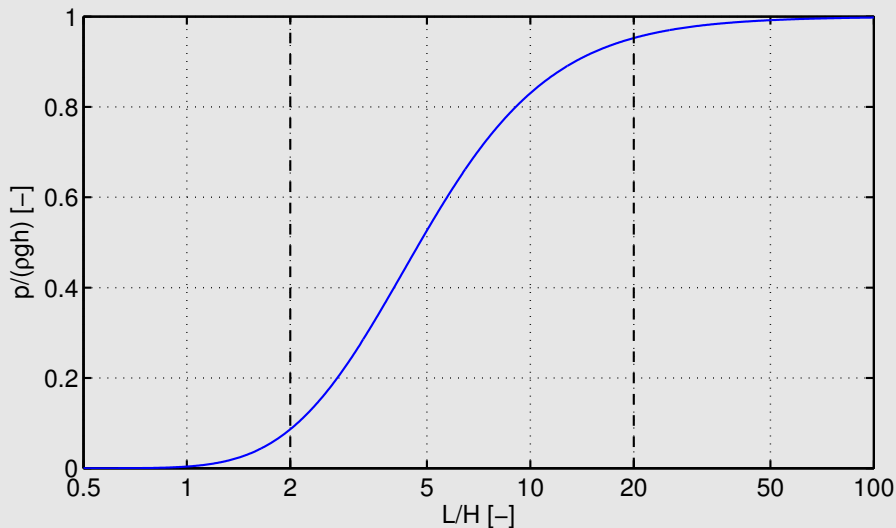
$$= \frac{h}{\sinh\left(\frac{H}{d}\right)} \rho e^{i\omega(t-s_1x_1)} \frac{g}{d} \tanh\left(\frac{H}{d}\right) d \cosh\left(\frac{x_3+H}{d}\right) \quad (38)$$

$$= \rho gh e^{i\omega(t-s_1x_1)} \frac{\cosh\left(\frac{x_3+H}{d}\right)}{\cosh\left(\frac{H}{d}\right)} \quad (39)$$

The Fluid Pressure



Variation in Pressure at the Ocean Floor



Variation in Pressure at the Ocean Floor

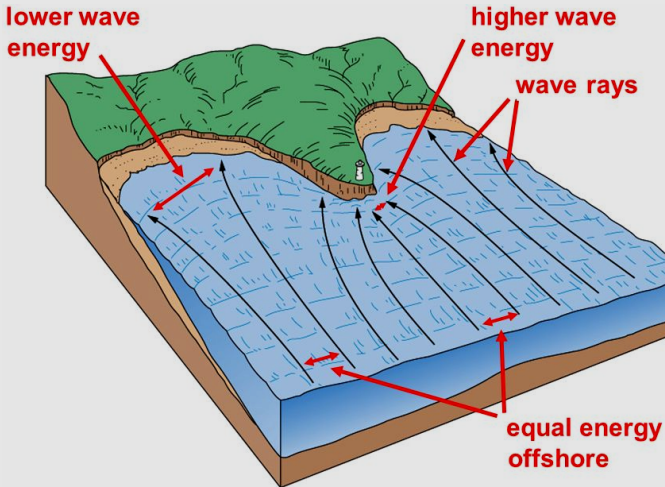
Deep water regime ($\frac{L}{H} \leq 2$): < 10 % of the near-surface variation at the ocean floor.

Shallow water regime ($\frac{L}{H} \geq 20$): > 95 % of the near-surface variation at the ocean floor.



Most important component of tsunami warning systems beyond earthquake registration.

Wave Shoaling



Source: Carpenter, Ocean Waves

Ray Theory

Extensions towards the harmonic plane wave approach:

- Retarded time $\tau = t - \psi(x_1, x_2)$ instead of $\tau = t - s_1 x_1$ with a general phase function $\psi(x_1, x_2)$



Propagation in direction of $\nabla\psi(x_1, x_2)$ with local slowness $|\nabla\psi(x_1, x_2)|$

- Spatially variable wave height $h(x_1, x_2)$
- Vertical particle displacement in analogy to Eq. 29,

$$u_3(\vec{x}, t) = h e^{i\omega(t-\psi(x_1, x_2))} \frac{\sinh\left(\frac{x_3+H}{d}\right)}{\sinh\left(\frac{H}{d}\right)} \quad (40)$$

so that $u_3 = 0$ at the ocean floor.

Ray Theory

Calculations in analogy the eikonal equation for seismic waves.



Terms $\sim \omega^2$:

- Horizontal particle displacement only in direction of propagation
- Velocity of propagation according to Eq. 35

Terms $\sim \omega$:

$$\operatorname{div}(\vec{q}) = 0 \quad (41)$$

with the energy flux density

$$\vec{q} = \frac{1}{2} \rho g h^2 \vec{v} \quad (42)$$